

Dynamic-Mode Decomposition of Geostrophically Balanced Motions from SWOT Cal/Val in the separated Gulf Stream

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Key Points:

- Dynamic-mode decomposition (DMD) is applied for the first time to sea-surface height (SSH) fields.
- DMD extracts the sub-inertial signals from SSH fields that have an imprint of internal gravity waves (IGWs).
- Slowly varying DMD spatial modes can be used to isolate geostrophically balanced motions.

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Abstract

The decomposition of oceanic flow into its geostrophically balanced and unbalanced motions carries theoretical and practical significance for the oceanographic community. These two motions have distinct dynamical characteristics and affect the transport of tracers differently from one another. The launch of the Surface Water and Ocean Topography (SWOT) satellite provides a prime opportunity to diagnose the surface balanced and unbalanced motions on a global scale at an unprecedented spatial resolution. Here, we apply dynamic-mode decomposition (DMD), a linear-algebraic data-driven method, to tidally-forced idealized and realistic numerical simulations at submesoscale-permitting resolution and one-day-repeat SWOT observations of sea-surface height (SSH) in the Gulf Stream downstream of Cape Hatteras, a region commonly referred to as the separated Gulf Stream. DMD is able to separate out the spatial modes associated with sub-inertial periods from super-inertial periods. The sub-inertial modes of DMD can be used to extract geostrophically balanced motions from SSH fields, which have an imprint of internal gravity waves, so long as the data extends long enough in time. We utilize the statistical relation between relative vorticity and strain rate as the metric to gauge the extraction of geostrophy.

Plain Language Summary

Observations of the global ocean surface are now done routinely by satellites. One of the key variables in describing the oceanic state is sea-surface height (SSH), i.e., elevations of the sea surface. For those who enjoy marine sports, it is well appreciated that the ocean surface is teeming with waves and currents. Similar to the density interface between the ocean and atmosphere, there are waves beneath the surface at density interfaces within the ocean. Waves at the ocean surface are called surface waves and in the interior are called internal waves. Both surface- and internal-wave signals imprint onto SSH. In order to extract information on oceanic currents (e.g., flow direction and speed) from SSH, it is necessary to remove the signal of surface and internal waves since waves and currents are generally not physically related to each other on the same scales in space and time. Namely, waves tend to propagate much faster and have smaller spatial scales than the currents. Here, we implement a method based on linear algebra, which is able to capture the slowly varying residual signals from the waves.

56 **1 Introduction**

57 With the launch of the Surface Water and Ocean Topography (SWOT) satellite,
 58 there is great interest within the oceanographic community to extract surface velocity
 59 information from the new altimetry observations with $\mathcal{O}(1\text{ km})$ spatial resolution (Dibarboure
 60 et al., 2024; J. Wang et al., 2024). The fact, however, that the observed altimetry is a
 61 superposed signal of geostrophic turbulence and waves complicates the problem (e.g. Rich-
 62 man et al., 2012; Savage et al., 2017; Le Guillou et al., 2023; Xiao et al., 2023; Maingonnat
 63 et al., 2024). While geostrophy is one of the most simple and practical balances that re-
 64 lates sea-surface height (SSH) gradients to velocity, horizontal gradients of unfiltered SSH
 65 observations are contaminated by high-frequency balanced motions and unbalanced mo-
 66 tions such as flows with Rossby numbers on the order of unity and larger and internal
 67 gravity wave (IGW) signals (Torres et al., 2018; McWilliams, 2019, 2021; Cao et al., 2023).

68 One work around has been to exploit the fact that submesoscale dynamics and IGWs
 69 are associated with smaller spatial scales and shorter time scales than geostrophic ed-
 70 dies. Namely, filtering the SSH and/or momentum fields by band-pass filters in the wavenum-
 71 ber and frequency domain (C. Wang et al., 2023a; Jones et al., 2023; Bakhoday Paskyabi,
 72 2024). A limitation of this approach is that Fourier transforms require the data to be
 73 periodic and to not have any gaps. Another popular method for modal decomposition,
 74 empirical orthogonal function (EOF), is excellent at extracting spatial modes of the data
 75 but decouples the space-time information (Uchida et al., 2021); the EOF spatial modes
 76 are unaware of the temporal phase information. Additionally, decomposition methods
 77 based only on spatial information do not remove IGWs that have wavelengths compa-
 78 rable to the local Rossby radii of deformation (Cao et al., 2021). Lagrangian filtering,
 79 on the other hand, requires direct knowledge of the momentum fields themselves (Shakespeare
 80 et al., 2021; C. Wang et al., 2023b; Jones et al., 2023; Baker et al., 2024; Minz et al., 2024),
 81 which SSH observations do not directly provide.

82 Here, we implement a relatively novel data-driven method coined as dynamic-mode
 83 decomposition (DMD) that decomposes the data into spatial modes while retaining the
 84 phase (i.e., growing, decaying and/or oscillating in time) information associated with each
 85 mode (Kutz et al., 2016). Conceptually, it can be thought of as applying the band-pass
 86 filter in the real space-time domain (instead of the wavenumber-frequency domain); or,
 87 it can be thought of as EOF spatial modes associated with temporal phase information.

88 DMDs have been widely adopted in the broader field of fluid mechanics (Schmid, 2022;
 89 Baddoo et al., 2023; S. L. Brunton & Kutz, 2024), plasma physics and geomagnetics (Chi-
 90 Durán & Buffett, 2023; Kutz et al., 2024), neuroscience (B. W. Brunton et al., 2016),
 91 and epidemiology (Proctor & Eckhoff, 2015). In the context of SWOT, our quest is to
 92 decompose the slowly varying geostrophic dynamics in first-order balance with Earth’s
 93 rotation and vertical stratification from the fast unbalanced motions associated with sub-
 94 mesoscale dynamics and IGWs given the observed SSH fields. As we shall see, DMD is
 95 capable of separating out the slow (sub-inertial) component in SSH without the require-
 96 ment of periodicity and will allow us to diagnose geostrophy from it. We will demonstrate
 97 that this method is an effective approach to isolate geostrophic motions in idealized and
 98 realistic high-resolution ocean simulations with IGWs, and one-day-repeat SWOT tracks.

99 The paper is organized in a way that demonstrates the application of multi-resolution
 100 coherent spatiotemporal scale separation (mrCOSTS), a variant of DMD, to flows from
 101 idealized configurations to increasing levels of complexity and realism. We briefly intro-
 102 duce the math behind mrCOSTS and the SSH dataset from idealized and realistic tidally-
 103 forced submesoscale-permitting simulations in the section below. We present our results
 104 in Section 3. Section 4 ends with a Discussion on Level 3 SWOT Calibration and Val-
 105 idation (Cal/Val) data (Dibarboure et al., 2024).

106 2 Method and Data

107 2.1 Multi-Resolution Coherent Spatiotemporal Scale Separation

108 At the basic level, dynamic-mode decomposition (DMD) is a method that seeks a
 109 locally linear dynamical system (Kutz et al., 2016)

$$\frac{d}{dt}\boldsymbol{\eta} = \mathcal{A}\boldsymbol{\eta}, \quad (1)$$

110 where \mathcal{A} is a linear operator and approximately encapsulates all physical processes re-
 111 sponsible for the system to step forward in time. In discrete form, this can be recasted
 112 as

$$\boldsymbol{\eta}_n = \mathbf{A}\boldsymbol{\eta}_{n-1} = \mathbf{A}^n\boldsymbol{\eta}_0, \quad (2)$$

113 where $\mathbf{A} = \exp(\mathcal{A}\Delta t)$ and $n = 1, 2, \dots$ is the time step. Δt is the time between the
 114 time steps when discretizing (1). The goal of DMD is to determine \mathbf{A} so that the solu-
 115 tion to (2) can be expressed by the eigenvalues λ and eigenvectors $\boldsymbol{\psi}$ of the discrete-time

116 map \mathbf{A}

$$\boldsymbol{\eta}_n = \sum_{j=1}^r \psi_j \lambda_j^n \mathbf{b}_j, \quad (3)$$

117 where \mathbf{b}_j are the coordinates of the initial state $\boldsymbol{\eta}_0$ in the eigenvector basis, and r is the
 118 rank of singular-value decomposition (SVD) of \mathbf{A} . Equation (2) can be expanded with-
 119 out a loss of generality as

$$\mathbf{H} = \begin{bmatrix} | & | & \cdots & | \\ \boldsymbol{\eta}_0 & \boldsymbol{\eta}_1 & \cdots & \boldsymbol{\eta}_{n-1} \\ | & | & & | \end{bmatrix}, \quad (4a)$$

$$\mathbf{H}' = \begin{bmatrix} | & | & & | \\ \boldsymbol{\eta}_1 & \boldsymbol{\eta}_2 & \cdots & \boldsymbol{\eta}_n \\ | & | & & | \end{bmatrix}, \quad (4b)$$

120 where \mathbf{H} and \mathbf{H}' are shifted by one time step. The DMD algorithm produces a low-rank
 121 eigen decomposition (3) of matrix \mathbf{A} that optimally minimizes the Frobenius norm (Askham
 122 & Kutz, 2018)

$$\|\mathbf{H}' - \mathbf{A}\mathbf{H}\|_F. \quad (5)$$

123 The approximation (5) arises from fitting a locally linear system (1) to a system that
 124 is in fact non-linear. By rewriting $\omega_j = \ln(\lambda_j)/\Delta t$, the approximate solution for all fu-
 125 ture times can be predicted as

$$\boldsymbol{\eta}(t, \mathbf{x}) \approx \sum_{j=1}^r \psi_j(\mathbf{x}) \exp(\omega_j t) \mathbf{b}_j = \boldsymbol{\Psi} \exp(\boldsymbol{\Omega} t) \mathbf{b}. \quad (6)$$

126 The real part of ω_j , $\text{Re}[\omega_j]$ gives growing or decaying modes in time while the imaginary
 127 part $\text{Im}[\omega_j]$ corresponds to oscillating modes. Equation (6) may look similar to EOFs,
 128 viz.

$$\boldsymbol{\eta}(t, \mathbf{x}) \approx \sum_{j=1}^M \mathbf{a}_j(t) \phi_j(\mathbf{x}), \quad (7)$$

129 where \mathbf{a} is the principle components and ϕ is the EOF spatial modes (cf. Uchida et al.,
 130 2021). The difference is that the spatial modes are decoupled from the temporal phase
 131 information in EOFs while the two are interlinked in DMDs.

132 In practice, we employ a variant of DMD, viz. multi-resolution coherent spatiotem-
 133 poral scale separation (mrCOSTS), a method which was originally proposed by Dylewsky
 134 et al. (2019) and advanced by Lapo et al. (2025), to deal with datasets comprising of multi-
 135 scale non-linear dynamics by iteratively applying DMD over the entire dataset. For each
 136 decomposition level a sliding window of fixed length in time is applied to the data and

137 a DMD model is fit, resulting in a collection of DMD models for each window. These
 138 DMD models are categorized into a poorly resolved low-frequency component and bet-
 139 ter resolved high-frequency components, called the local-scale separation. The high- and
 140 low-frequencies discovered are in reference to the window length. The low-frequency com-
 141 ponent is used as input to the next decomposition level with a larger window size. Namely,
 142 the highest frequency components are extracted at each level. This local-scale process
 143 is iterated over the the number of *a priori* decomposition levels prescribed by the user.
 144 Upon completion of the local scale a global-scale separation is performed, which captures
 145 leaked frequency components between decomposition levels. The global-scale separation
 146 is achieved by applying the *k*-means clustering to the temporal dynamics of the collec-
 147 tion of DMD models (Pedregosa et al., 2011). Namely, we collect the bands across all
 148 decomposition windows for the global-scale separation to generate the bands (see also
 149 Lapo et al., 2025, their Fig. 1).

150 Due to subtracting out the mean within each window, mrCOSTS is especially amenable
 151 to diagnosing fluid flows as the decomposition approximates the high-frequency fluctu-
 152 ations of a Reynolds' averaged flow at each decomposition level. The resulting decom-
 153 position identifies discrete bands of coherent spatiotemporal modes. The scale-separated
 154 bands are denoted $\mathcal{G}_1, \mathcal{G}_2, \dots, \mathcal{G}_p, \dots, \mathcal{G}_P$ where the subscript p indexes the scale-separated
 155 bands and P is the total number of bands. Each band can then be used to reconstruct
 156 the contribution of \mathcal{G}_p to the original data, $\check{\boldsymbol{\eta}}_p(t)$,

$$\check{\boldsymbol{\eta}}_p(t, \mathbf{x}) = \sum_{k=1}^N \sum_{(j,\ell) \in \mathcal{G}_p} \boldsymbol{\psi}_{j,\ell}^k(\mathbf{x}) \exp(\omega_{j,\ell}^k t) \mathbf{b}_{j,\ell}^k. \quad (8)$$

157 The subscript ℓ denotes the decomposition level and j to index the DMD eigenvalue ω
 158 and eigenvector $\boldsymbol{\psi}$ pairs up to rank r specific to the ℓ^{th} level. The superscript k is used
 159 to index the data windows so that snapshots belonging to the k^{th} window are approx-
 160 imated by the decomposition. We use $\check{\boldsymbol{\eta}}_p(t)$ to indicate an approximation of the origi-
 161 nal input signal per band $\boldsymbol{\eta}_p(t)$. Summing over a subset of the bands, p , allows one to
 162 reconstruct a slow and fast component of the data. The mrCOSTS reconstruction of the
 163 original data is achieved by summing up over all bands in addition to the background
 164 band, $\boldsymbol{\eta}(t) \approx \sum_p^P \check{\boldsymbol{\eta}}_p(t) + \check{\boldsymbol{\eta}}_b(t)$. The background (lowest-frequency) band, $\check{\boldsymbol{\eta}}_b(t)$, is the
 165 left-over low-frequency component after finishing recursively applying mrCOSTS as the
 166 high-frequency component gets extracted at each level.

167 MrCOSTS can provide a robust scale separation for a range of hyperparameters,
 168 often requiring little-to-no tuning. The most relevant hyperparameters are the length
 169 of the window used at each decomposition level, the SVD rank of the DMD fit at each
 170 level, any constraints on the eigenvalue solutions and eigenvalues themselves. We refer
 171 the reader to Lapo et al. (2025) and Ichinaga et al. (2024, their online tutorial [https://](https://github.com/PyDMD/PyDMD/tree/master/tutorials/tutorial120)
 172 github.com/PyDMD/PyDMD/tree/master/tutorials/tutorial120) for further details re-
 173 garding the implementation and user guide on mrCOSTS.

174 2.2 Idealized Wave-Vortex Simulation

175 As was demonstrated by Early et al. (2021), an unambiguous decomposition be-
 176 tween linear waves and geostrophic motions in variable stratification can be made un-
 177 der flat-bottom boundary conditions. The eigenmodes from the decomposition (Early,
 178 Hernández-Dueñas, et al., 2024) form a spectral basis for the wave-vortex model (Early,
 179 Avila, et al., 2024), which then solves the equations of motion for a doubly-periodic ro-
 180 tating non-hydrostatic Boussinesq fluid with arbitrary stratification. At each instant in
 181 time the complete state of the fluid is decomposed into geostrophic and wave modes, while
 182 the nonlinear time steps flux energy between modes. Although other methodologies for
 183 separating waves and geostrophic motions exist, they are either restricted to constant
 184 stratification and shallow-water systems (Chouksey et al., 2023, & references therein),
 185 or use a temporal filter that depends on the linear dispersion relation for waves (Lelong
 186 et al., 2020; Shakespeare et al., 2021). In contrast, the wave-vortex decomposition is fun-
 187 damentally an inversion of quasi-geostrophic potential vorticity, and separates the flow
 188 without any assumptions about its temporal evolution. The advantage to this approach
 189 is that the wave-vortex decomposition and DMD are using very different information about
 190 the flow, and thus make comparisons all the more meaningful.

191 For the simulation considered here, the ocean was spun up by imposing bottom fric-
 192 tion and continually relaxing to a weak, low-mode geostrophic flow, which goes baroclin-
 193 ically unstable; this injects energy and enstrophy into the system, which is then removed
 194 by the bottom friction and small-scale damping. In 2000 days, the flow reaches steady-
 195 state, resulting in a single pair of dipolar geostrophic eddies, after which inertial oscil-
 196 lations and a narrow band of IGWs with semidiurnal frequency were prescribed as a forc-
 197 ing. Interactions with the mesoscale eddy field cause a robust IGW field to emerge (Lelong
 198 et al., 2020) which reaches a steady-state IGW field within 10 days that resembles a Garrett-

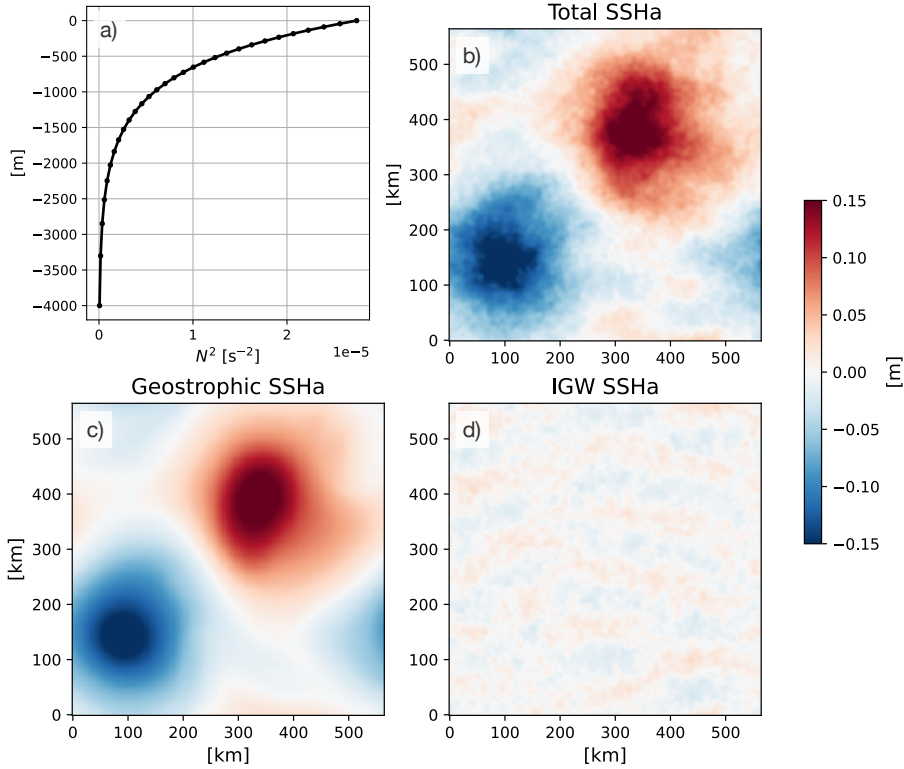


Figure 1. Buoyancy frequency (N^2) with an exponential vertical profile (a) and snapshots of the total SSHa and its geostrophic and IGW components at an arbitrary time step (i.e., the 200th time step after the flow is spun up) from the doubly-periodic wave-vortex simulation (b-d). N^2 is kept stationary throughout the simulation.

199 Munk spectrum (Garrett & Munk, 1975) with an amplitude close to the level of SWOT
 200 instrumental noise. The inertial and semi-diurnal forcing is maintained during the 70-
 201 day analysis period where geostrophic and IGW fields evolve through nonlinear inter-
 202 actions. This configuration is ideal to test mrCOSTS as we will know the exact initial
 203 frequencies of the waves. Furthermore, because the wave-vortex model exactly isolates
 204 the geostrophic component at each time step, this will be treated as our target for mr-
 205 COSTS to extract from the total SSH anomaly (SSHa). The prescribed vertical strat-
 206 ification and snapshots of the spun-up SSHa of the geostrophic and IGW component at
 207 an arbitrary time step is documented in Fig. 1. An animation of the spun-up fields of
 208 relative vorticity normalized by the Coriolis frequency is provided in the Supporting In-
 209 formation (dmd-movie.mp4) where the local Rossby numbers are small ($\text{Ro} = \zeta/f \sim$
 210 $\mathcal{O}(0.1)$).

211 Given that we know *a priori* that the flow consists of a single pair of geostrophic
 212 eddies and IGWs with distinct frequencies from each other, we applied mrCOSTS in two
 213 levels with the window lengths of [1, 8] days respectively; this splits the SSH anomaly (SSHa)
 214 fields into high- and low-frequency DMD modes about the window lengths during each
 215 iteration. In other words, the number of iterations for the local-scale separation here was
 216 prescribed as two ($N = 2$ in (8)). The first window length was chosen to be diurnal and
 217 the second window length is the characteristic time scale of geostrophic eddies (cf. Tor-
 218 res et al., 2018, their Fig. 3). The model outputs were saved every 30 minutes but hourly
 219 resolution was used to construct \mathbf{H} and \mathbf{H}' . Conceptually, for a one-day (24-hour) and
 220 eight-day window, mrCOSTS fits 24 and 192 data points respectively in time for data
 221 with hourly resolution ($\Delta t = 1$ hour). Namely, the number of data points per window
 222 depends on the temporal resolution of the data used. The window is then slid in time
 223 to go through the entire dataset in a manner similar to how one would take the running
 224 mean. The ranks of SVD were set as [8, 18], which need to be smaller than the number
 225 of data points within each window, i.e., [24, 192] respectively. Increasing the ranks gen-
 226 erally leads to mrCOSTS finding more modes, $\psi_{j,\ell}^k$, but given the simplicity of the flow,
 227 we have kept it minimalistic.

228 Figure 2a shows the probability density function (PDF) of frequencies associated
 229 with spatially coherent modes discovered by mrCOSTS and the frequency spectrum of
 230 SSHa over the duration of 72 days; periodograms were taken every ~ 150 km and then
 231 spatially averaged to construct the spectrum. We see that the SSHa fields contain a sig-
 232 nal of IGWs with a peak around the diurnal and semidiurnal frequencies.

233 **2.3 Tidally-Forced Submesoscale-Permitting North Atlantic Simulation**

234 We take the hourly SSHa snapshot outputs from an atmospherically and tidally
 235 forced North Atlantic simulation at $1/50^\circ$ ($\mathcal{O}(2\text{ km})$) resolution using the HYbrid Co-
 236 ordinate Ocean Model (HYCOM50; Xu et al., 2022); data of subdomains are publicly
 237 available via the Open Storage Network, a cloud storage service operated by the National
 238 Science Foundation (NSF; Uchida et al., 2022). Bathymetry data was taken from the 15-
 239 second GEBCO 2019 global dataset and the modeled domain covers the North Atlantic
 240 from 28°S to 80°N . HYCOM50 was spun up for 15 years from the U.S. Navy’s Global
 241 Ocean Climatological (GDEM) state of rest and forced with the monthly climatologi-
 242 cal European Centre for Medium-Range Weather Forecasts (ECMWF) Reanalysis ERA-

243 40. Additionally, three-hourly wind anomalies from the Navy Operational Global Atmo-
 244 spheric Prediction System (NOGAPS) and the Climate Forecast System Reanalysis (CFSR)
 245 for the year 2003 (a neutral year in term of large-scale North Atlantic Oscillation pat-
 246 tern) were prescribed with absolute wind stress. Eight tidal constituents were included
 247 (K_1 , O_1 , P_1 , Q_1 , M_2 , S_2 , N_2 , and K_2). The three months of August–October (ASO) in
 248 year 19 is used in our analyses below. Further details on the model configuration can be
 249 found in Chassignet and Xu (2017) and Xu et al. (2022).

250 In constructing \mathbf{H} and \mathbf{H}' , we sub-sampled the SSHa fields every three and 12 hours
 251 in the separated Gulf Stream ($\Delta t = 3$ and 12 hours; Fig. 4a), a region partially over-
 252 lapping with a SWOT crossover (Figs. 4a and 6). The temporal sub-sampling mimics
 253 observations where high resolution in time is not obtainable. The spatial mean was re-
 254 moved from each snapshot and the fields were further spatially smoothed by applying
 255 a Gaussian filter with the standard deviation of 10 km using the `gcm-filters` Python
 256 package (Grooms et al., 2021); we do not expect perturbations on scales smaller than
 257 this to be in geostrophic balance (Pedlosky, 1984; Vallis, 2006) and some spatial filter-
 258 ing is justified to compensate for the lack of temporal resolution. The latitude-longitude
 259 dimensions were flattened into a one-dimensional array so as to feed mrCOSTS two-dimensional
 260 fields in space-time.

261 We applied mrCOSTS in six levels ($N = 6$) with each level splitting the SSH anomaly
 262 (SSHa) fields into high- and low-frequency modes about the window lengths of [1, 2, 3, 4, 8, 30] days
 263 respectively for the three-hourly case. The first four window lengths were chosen to be
 264 close to tidal periods and their harmonics, the fifth window length is the characteristic
 265 time scale of geostrophic eddies, and the sixth window length has a monthly time scale.
 266 When the data were sub-sampled 12 hourly, we applied mrCOSTS in four levels ($N =$
 267 4) with each window length corresponding to [3, 4, 8, 30] days. The ranks of SVD were
 268 set as [6, 6, 6, 8, 8, 10] in the three-hourly case for each level and [4, 4, 6, 10] in the 12-hourly
 269 case. Note that shortening the window lengths will potentially lead to discovering bands
 270 with higher frequencies but they must be long enough to allow for the least-squares fit
 271 (5) and SVD to converge.

272 Figure 2b shows the PDF of frequencies discovered by mrCOSTS and the frequency
 273 spectrum of SSHa over the three months of August - October; periodograms were taken
 274 every ~ 100 km and then spatially averaged in constructing the spectrum. We see that

275 the SSHa fields contain a strong signal of internal tides with peaks around the diurnal
 276 and semidiurnal frequencies. Focusing on the PDF, one notices that mrCOSTS bands,
 277 \mathcal{G}_p , around higher-order tidal harmonics are missing. While further tuning of the param-
 278 eters (e.g., number of decomposition levels, window length prescribed to each level, rank
 279 of SVD, etc.) may improve the discovery of super-inertial frequencies, given that our in-
 280 terest here is in extracting the sub-inertial geostrophic dynamics, we have settled with
 281 the parameter settings described above. It is possible that Doppler shift could lead to
 282 a shift in frequencies, and thus part of the balanced dynamics could be associated with
 283 super-inertial frequencies (Chouksey et al., 2018). Our focus, however, remains on the
 284 sub-inertial signal as the component in first-order balance with Earth’s rotation and ver-
 285 tical stratification.

286 3 Results

287 The mrCOSTS parameters used for each experiment are summarized in Table 1.

Table 1. MrCOSTS parameters used for each experiment

Experiment	Levels	Window lengths	SVD ranks	Bands comprising the slow mode
Wave-vortex	$N = 2$	[1, 8] days	[8, 18]	$p = [0, 1]$
–	$N = 3$	[0.5, 1, 2, 16] days	[4, 8, 10, 18]	$p = [0, 1, 2]$
HYCOM50 (3 hourly)	$N = 6$	[1, 2, 3, 4, 8, 30] days	[6, 6, 6, 8, 8, 10]	$p = [0, 1, 2, 3, 4, 5, 6]$
– (12 hourly)	$N = 4$	[3, 4, 8, 30] days	[4, 4, 6, 10]	$p = [0, 1, 2, 3, 4, 5, 6, 7]$
– (24 hourly)	$N = 4$	[9, 10, 11, 30] days	[4, 4, 6, 10]	$p = [0, 1, 2, 3]$
– (24 hourly, JASON)	$N = 5$	[9, 10, 11, 30, 90] days	[4, 4, 6, 10, 18]	$p = [0, 1, 2, 3, 4]$
SWOT Cal/Val	$N = 4$	[9, 10, 11, 30] days	[4, 4, 6, 10]	$p = [0, 1, 2, 3, 4, 5]$

288

289 3.1 Idealized Wave-Vortex Simulation

290 As a proof of concept, we start by demonstrating the spatial maps of SSHa recon-
 291 struction. The mrCOSTS reconstruction of the sub-inertial (slow) component of SSHa

292 from the wave-vortex simulation is shown in Fig. 3b. In total, mrCOSTS discovered 12
 293 bands ($p = 0, 1, \dots, 11$) based on the convergence of SVD (the frequency associated
 294 with each band is given in Fig. 2a); this is similar to EOF where it yields a finite num-
 295 ber of modes (M in (7)) or discrete spectral decomposition based on fast-Fourier trans-
 296 form where the number of Fourier modes is determined by the data resolution and Nyquist
 297 frequency/wavenumber. Given the 12 bands in total, the decision on what to select as
 298 part of the ‘slow’ component becomes somewhat subjective. Here, we chose the slow com-
 299 ponent as the net sum of the background band and first two bands ($p = 0, 1$), which
 300 are associated with periods longer than a day. Figure 3a shows the time series of the spa-
 301 tial correlation between the geostrophic component and mrCOSTS slow component; the
 302 spatial correlation is persistently higher than 0.999. The correlation was computed as:

$$\rho = \frac{\mathcal{E}[(\eta - \bar{\eta})(\sum_p \check{\eta}_p + \check{\eta}_b - \overline{\sum_p \check{\eta}_p + \check{\eta}_b})]}{\sigma_\eta \sigma_{\check{\eta}}}, \quad (9)$$

303 where $\mathcal{E}[\cdot]$ is the expectation value, $\overline{(\cdot)}$ is the spatial mean, and σ_μ and $\sigma_{\check{\mu}}$ are the spa-
 304 tial standard deviations of the geostrophic component and mrCOSTS slow component
 305 respectively. The difference between Figs. 1c and 3b, c is hardly detectable by the naked
 306 eye. This indicates that the slow component of mrCOSTS can be used to diagnose sub-
 307 inertial motions in geostrophic balance. The slightly lower correlation towards the be-
 308 ginning and ending of the time series is attributed to $\check{\eta}_p(t)$ having the largest errors at
 309 the edges of the time domain due to edge effects analogous to the cone-of-influence (COI)
 310 in wavelet analysis (Torrence & Compo, 1998; De Moortel et al., 2004; Lapo et al., 2025).

311
 312 Now, we can play the game where we assume that we had no prior knowledge of
 313 the flow. Namely, a case where, from eye inspection, we can tell that the flow consists
 314 of eddies and waves (Fig. 1b) but do not know the exact frequencies of the dynamics.
 315 Based on the frequency spectrum (Fig. 2a), we can make an educated guess that the flow
 316 has peaks about the semidiurnal and diurnal frequencies so we can prescribe the win-
 317 dows lengths as [0.5, 1, 2, 16] days. The corresponding SVD ranks were set as [4, 8, 10,
 318 18]. MrCOSTS found eight bands in total and we chose the first three bands with pe-
 319 riods longer than a day as part of the slow component (Table 1; Supporting Information
 320 Fig. S1a). We again find that mrCOSTS decomposes and reconstructs SSHa with small
 321 errors (on the order of 1%; Fig. 3a, c and e). The point of all this is that the mrCOSTS
 322 algorithm is highly versatile to the choice of parameters (as was noted in Section 2.1)

323 so long as the window lengths cover a reasonable range of distinct time scales for phe-
 324 nomena consisting of multi-scale dynamics (Lapo et al., 2025).

325 **3.2 Tidally-Forced Submesoscale-Permitting North Atlantic Simulation**

326 Encouraged by the success from the idealized case, we show the modeled SSHa field
 327 in HYCOM50 in Fig. 4a and the mrCOSTS reconstruction of its sub-inertial (slow) com-
 328 ponent. The slow component was chosen to be the net sum of the background band and
 329 bands 0-7, which have periods longer than 2 days (i.e., the vertical black dashed lines
 330 in Fig. 2b associated with frequencies lower than 5×10^{-1} cpd). We see that slowest (back-
 331 ground) band already captures the large-scale features of the separated Gulf Stream and
 332 a cold-core eddy (Fig. 4b). The addition of bands up to seven further improves the re-
 333 construction when the SSHa fields are fed every three hours to construct \mathbf{H} and \mathbf{H}' (Fig. 4c, d).
 334 This is corroborated by the spatial correlation shown in Fig. 4g where outside of COI,
 335 the correlation is always higher than 0.99. We also find that the performance of mrCOSTS
 336 remains relatively insensitive to temporal sub-sampling. This is highlighted by the spa-
 337 tial maps and spatial correlation where the SSHa fields were given every 12 hours (Fig. 4e-
 338 g). The first eight bands out of the 10 were summed up to obtain the slow component
 339 for the 12-hourly case (Table 1; Fig. S1b in Supporting Information).

340 Given the extraction of the slow component of SSHa evolution, we can diagnose
 341 geostrophy from the fields

$$fu = -g\eta_y, \quad fv = g\eta_x, \quad (10)$$

342 and from it, relative vorticity $\zeta = v_x - u_y$ and strain rate $|\alpha| = \sqrt{(u_x - v_y)^2 + (v_x + u_y)^2}$.
 343 Since relative vorticity and strain rates are second-order derivative terms of SSHa, they
 344 will highlight the small-scale features (or the lack thereof; Shcherbina et al., 2013; Bal-
 345 wada et al., 2021; Jones et al., 2023). When the spatially-smoothed instantaneous snap-
 346 shot outputs of SSHa fields are used, the imprint of IGWs contaminate the estimates of
 347 geostrophic relative vorticity (Fig. 5a). This is also indicated in the joint PDF of rela-
 348 tive vorticity and strain rates where there is an anomalously high likelihood of values with
 349 large amplitude and negative values in relative vorticity (Fig. 5e); geostrophy is only ex-
 350 pected to hold under small Rossby numbers (Vallis, 2006). The waves can be filtered out
 351 by taking daily averages of the hourly SSHa field (Fig. 5b, f). This gives us a reference

352 for geostrophic eddies, but we are interested in cases where hourly temporal resolution
 353 is not available at hand.

354 Figure 5c and d document the relative vorticity fields from the slow component ex-
 355 tracted by mrCOSTS. First thing to note is that the mrCOSTS bands are smooth enough
 356 to permit second-order spatial derivatives. There is a large attenuation in the signal from
 357 IGWs with its performance being better when SSHa fields are given every three hours
 358 compared to 12 hours. The same description also applies to both strain rate and hor-
 359 izontal divergence fields (Figs. S2 and S3). Nonetheless, both cases of sub-sampling cap-
 360 ture the joint PDF features of geostrophic eddies (Fig. 5g, h). The Rossby numbers on
 361 the order of unity ($Ro \sim \mathcal{O}(1)$) present in the mrCOSTS slow component and daily-
 362 averaged SSHa likely indicate that the eddies and meandering of the Gulf Stream anal-
 363 ysed here are in cyclogeostrophic balance (Fig. 5b, c; Hiron et al., 2021); these are sig-
 364 nals we want to retain in addition to geostrophy and mrCOSTS works surprisingly well
 365 in doing so.

366 4 Discussion and Conclusions

367 We end by discussing results on applying multi-resolution coherent spatiotempo-
 368 ral scale separation (mrCOSTS) to the one-day-repeat SWOT observations of SSHa ($\Delta t =$
 369 24 hours) during its Cal/Val phase (March 29–July 11, 2023). We have taken the Level
 370 3 (L3) KaRIn filtered product (Dibarboure et al., 2024) as our interest here is in extract-
 371 ing the first-order balance, geostrophy, from signals that include IGWs. The domain we
 372 use is between $30^\circ - 40^\circ\text{N}$ and $284^\circ - 288^\circ\text{E}$ for pass number nine situated across the sep-
 373 arated Gulf Stream path. Missing data and spacing between the swaths were linearly
 374 interpolated over and when data were missing from over 70% of the swaths, that day was
 375 dropped and temporally interpolated over between the day before and after via a lin-
 376 ear spline. The SWOT SSHa fields were further smoothed with a Gaussian filter with
 377 the standard deviation of 15 km. MrCOSTS was then applied with the period associated
 378 with each window length prescribed as [9, 10, 11, 30] days and rank of SVD as [4, 4, 6, 10]
 379 respectively. The slow component was defined as the sum of the background and first
 380 six bands (0-5) out of the 10 (Fig. 2c, Table 1).

381 When geostrophy (10) is applied directly to SWOT data, the relative vorticity and
 382 strain rate take fictitiously large values in magnitude despite the L3 product being some-

383 what smoothed via the de-noising process (Fig. 6e, g, i; Dibarboure et al., 2024); in hind-
 384 sight, the large magnitudes may have been expected as we are applying (10) to a field
 385 that includes signals of super-inertial balanced and unbalanced dynamics. The mrCOSTS
 386 slow component of SSHa, on the other hand, captures the large-scale feature and is much
 387 smoother than the SWOT data (Fig. 6a, b). The zonal geostrophic velocity from the mr-
 388 COSTS slow component captures the separated Gulf Stream about 37°N (Fig. 6d), and
 389 the fields of relative vorticity and strain rate become smoother and fall within the ac-
 390 ceptable range of magnitude ($\mathcal{O}(\text{Ro} \lesssim 1)$, Fig. 6f, h; Pedlosky, 1984, 2013; Vallis, 2006).
 391 Nonetheless, the joint PDF does not adequately capture the skewness towards positive
 392 relative vorticity values (Fig. 6j). As a reference, the joint PDF computed from daily-
 393 averaged 0.25° gridded AVISO data during April–June, 2023 is shown in Fig. 6k as the
 394 period that overlaps with the SWOT Cal/Val period; the skewness is only marginally
 395 captured and the magnitudes are much smaller, indicating that AVISO misses most of
 396 the frontal features. The spatial correlation between SWOT SSHa and its mrCOSTS re-
 397 construction is generally higher than 0.9 during the Cal/Val phase (Fig. 6l) but is worse
 398 than the case with the wave-vortex and HYCOM50 simulations. While there is a hint
 399 of mrCOSTS detecting the diurnal tidal signal (band nine in Fig. 2c), it is likely that
 400 the duration of three months with daily resolution (i.e., 102 data points in time) is not
 401 a sufficient amount of data to robustly estimate \mathbf{A} from the least-squares fit (5).

402 In order to test whether extending the duration of the data would improve the ex-
 403 traction of geostrophy, we examine the mrCOSTS reconstruction of HYCOM50 SSHa
 404 snapshot fields taken at daily intervals ($\Delta t = 24$ hours) when the duration to construct
 405 \mathbf{H} and \mathbf{H}' is taken over the three months of August to October (ASO), and five months
 406 of July to November (JASON). The window lengths were prescribed as [9, 10, 11, 30, 90] days
 407 and SVD ranks as [4, 4, 6, 10, 18] for the JASON case where 90 days corresponds to sea-
 408 sonal time scales. MrCOSTS was applied over four levels using the first four parame-
 409 ters for the ASO case (Table 1). MrCOSTS discovered seven bands in total for the JA-
 410 SON case ($p = 0, 1, \dots, 6$; Fig. S1c) and six bands for the ASO case ($p = 0, 1, \dots, 5$;
 411 Fig. S1d). We find that in both cases, mrCOSTS is able to reconstruct the HYCOM50
 412 SSHa fields relatively well; the spatial correlation outside of COI is higher than 0.98 (Fig. 7a).
 413 For the slow component, the sum of the first five bands were chosen for JASON and of
 414 the first four bands for ASO in addition to their respective background bands. We again
 415 diagnose the geostrophic relative vorticity and strain rates from the slow components and

416 the joint PDFs of the two are documented in Fig. 7b,c. Similar to SWOT (Fig. 6j), the
417 joint PDF from the ASO case does not present the skewness in relative vorticity. How-
418 ever, we find that the JASON case is able to recover the skewness and the joint PDF
419 becomes closer to Fig. 5f. While the SWOT Cal/Val phase is only available for three months,
420 this theoretical exercise of five months corroborates our hypothesis that the performance
421 of mrCOSTS depends somewhat on the number data points in time to fit (5). This sen-
422 sitivity to the volume and quality of data is not unique to DMD but rather universal to
423 data-driven methods (e.g. Budach et al., 2022; Chen et al., 2023; Smith et al., 2023; Mo-
424 jgani et al., 2024).

425 The goals of this paper were to introduce mrCOSTS, a variant of dynamic-mode
426 decomposition (DMD), to the oceanographic and earth science community. While machine-
427 learning methods have shown some promise in extracting the surface flow kinematics from
428 SSH (e.g., Sinha & Abernathey, 2021; H. Wang et al., 2022; Xiao et al., 2023; Gao et al.,
429 2024; Archambault et al., 2024; Cutolo et al., 2024; Fablet et al., 2024; Febvre et al., 2024;
430 Martin et al., 2024; Lyu et al., 2024), we have opted for DMD here due its interpretabil-
431 ity owing to it essentially being a combination of linear-algebraic operations. The fact
432 that DMD naturally decomposes the data into frequency components is also well suited
433 for disentangling geostrophically balanced motions from IGWs where the two tend to have
434 distinct characteristic time scales. We have showcased that by applying mrCOSTS to
435 modeled and observed SSHa, its slow bands are usable to diagnose geostrophy. In con-
436 trast to other DMD-based methods, mrCOSTS is able to robustly extract spatially co-
437 herent spatial modes (Lapo et al., 2025), which are smooth enough to permit spatial deriva-
438 tives. The need for scale-separation methods is wide spread in the general earth science
439 community; for example, it would be interesting to apply mrCOSTS to long-standing
440 problems such as quantifying orographic precipitation patterns (e.g., Buttafuoco et al.,
441 2011; Curio & Scherer, 2016; Y. Li et al., 2024) or discovering climate modes (e.g., New-
442 man et al., 2016; Dewar et al., 2022; Mishonov et al., 2024; G. Wang et al., 2024; Miyamoto
443 & Xie, 2024) and eddy parametrizations (L. Li et al., 2023).

444 Future work involves extending our analyses to other geographical regions, the 21-
445 day-repeat SWOT orbit where data will be available for longer periods than three months,
446 and to extract higher-order balances than geostrophy, viz., quasi- and semi-geostrophy.
447 The Southern Ocean may be an appealing region given the overlap amongst SWOT swaths
448 increases compared to lower latitudes. While we have purely focused on the geostroph-

449 ically balanced component of the flow in this study, it is true that information on the
450 unbalanced motion (e.g., IGWs) is also of significant value (Yadidya et al., 2024; Demat-
451 teis et al., 2024; Tchilibou et al., 2025). It is unclear to what extent DMDs can separate
452 out IGWs from submesoscale dynamics (or waves from turbulence in general; cf. Chávez-
453 Dorado et al., 2024), which tend to be associated with similar time scales, but will be
454 an avenue for further investigation.

455 **Open Research**

456 The wave-vortex model is available from Early, Avila, et al. (2024, [https://github](https://github.com/Energy-Pathways-Group/GLOceanKit)
457 [.com/Energy-Pathways-Group/GLOceanKit](https://github.com/Energy-Pathways-Group/GLOceanKit)) and the scripts for this particular simu-
458 lation are available at <https://github.com/JeffreyEarly/DMDEddySimulation>. The
459 HYCOM50 model outputs used in this paper are publicly available on the Open Stor-
460 age Network (OSN; <https://www.openstoragenetwork.org/>). Example Jupyter note-
461 books and Yaml file used to access the data are available on Github (Stern et al., 2022,
462 https://github.com/pangeo-data/swot_adac_ogcms). Jupyter notebooks used for anal-
463 yses will be shared with a DOI upon acceptance of the manuscript ([https://github.com/](https://github.com/roxyboy/GeostrophicDMD/tree/div)
464 [roxyboy/GeostrophicDMD/tree/div](https://github.com/roxyboy/GeostrophicDMD/tree/div)). Level 3 SWOT data (2-km, version 1.0) were ac-
465 cessed from Dibarboure et al. (2024, [https://www.aviso.altimetry.fr/en/data/products/](https://www.aviso.altimetry.fr/en/data/products/sea-surface-height-products/global/swot-13-ocean-products.html)
466 [sea-surface-height-products/global/swot-13-ocean-products.html](https://www.aviso.altimetry.fr/en/data/products/sea-surface-height-products/global/swot-13-ocean-products.html)). The altime-
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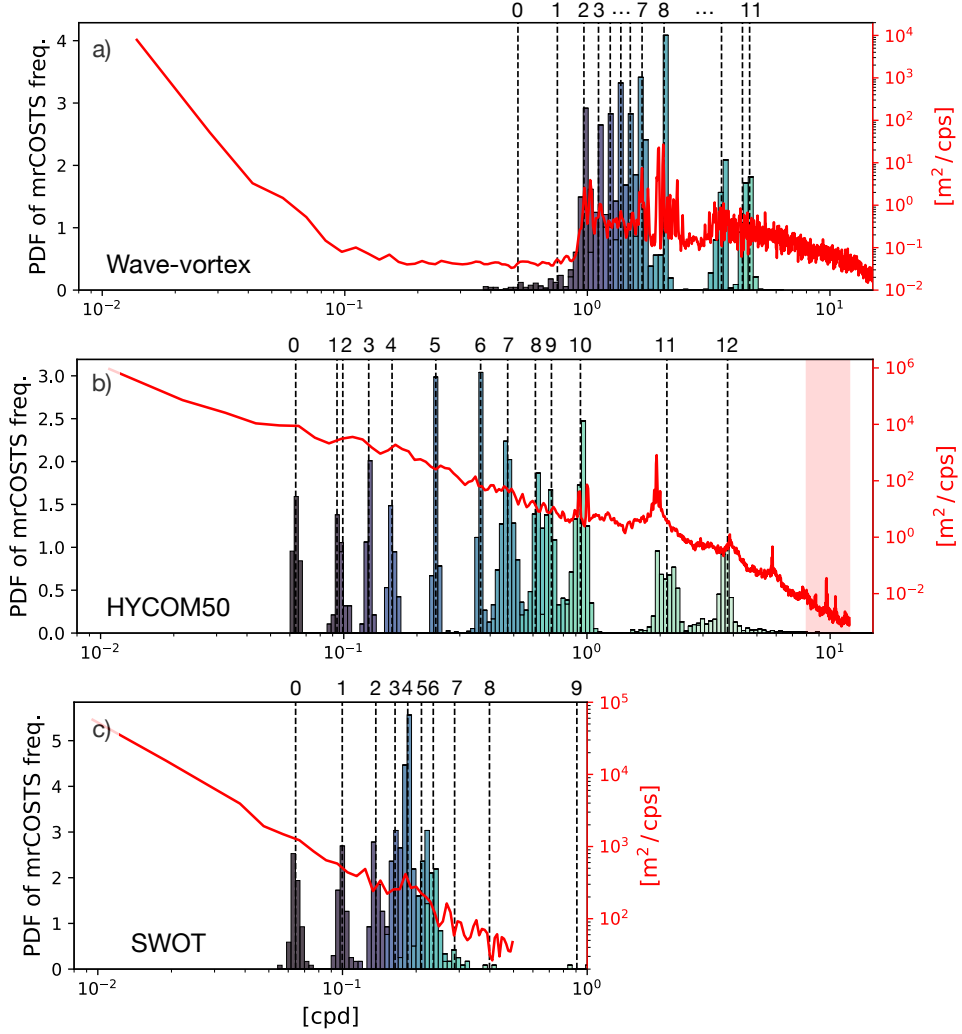


Figure 2. Probability density function (PDF) of the mrCOSTS-derived frequencies in cycles per day, $\text{Im}[\omega_{j,\ell}^k]/2\pi$. Wave-vortex SSHa fields were fed every hour to mrCOSTS and applied in two levels (a). The vertical black dashed lines indicate the frequencies each k -means clustering has grouped the mrCOSTS modes around for each band p . The top x axis shows the final mrCOSTS bands, viz. 12 bands in total ($p = 0, 1, \dots, 11$). The histogram is colored from light-to-dark shading corresponding to clusters from high-to-low frequency. Frequency spectrum of SSHa in red solid curve is plotted against the right y axis. HYCOM50 SSHa snapshot fields were fed every three hours (b). The red shading indicates the three-hour cutoff. MrCOSTS discovered 13 bands in total ($p = 0, 1, \dots, 12$). The same but for where instantaneous SWOT data were fed daily and mrCOSTS discovered 10 bands ($p = 0, 1, \dots, 9$; c). Periodograms were computed every 150 data points along track and 40 data points across track and then spatially averaged in constructing the SWOT frequency spectrum.

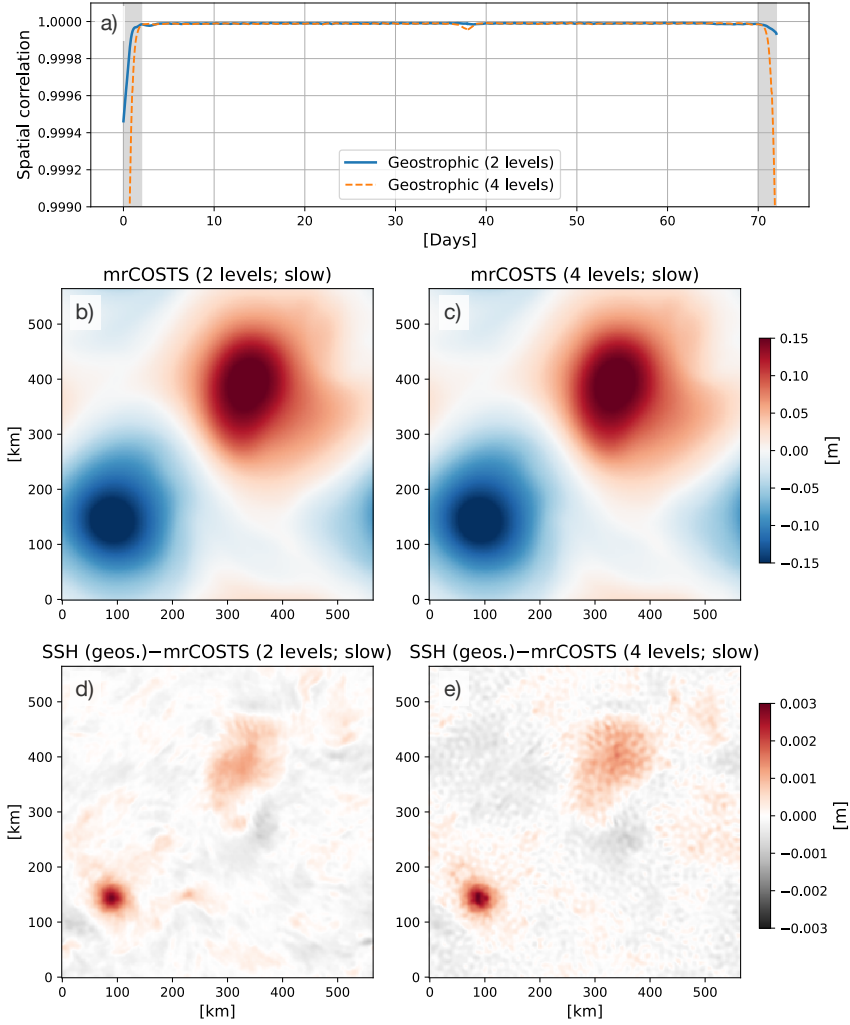


Figure 3. Time series of the spatial correlation between the geostrophic component from the wave-vortex SSH field and mrCOSTS reconstruction of the slow component (a). The black shading indicates the duration of COI and the x axis shows the number of days of model simulation. The spatial correlation for the $N = 2$ case is shown in solid blue and $N = 4$ case in dashed orange curves respectively. A snapshot of the mrCOSTS slow component on the same day as in Fig. 1c when mrCOSTS is applied over two levels ($N = 2$) (b) and four levels ($N = 4$) (c). The difference between the wave-vortex geostrophic component and mrCOSTS slow component (d,e).

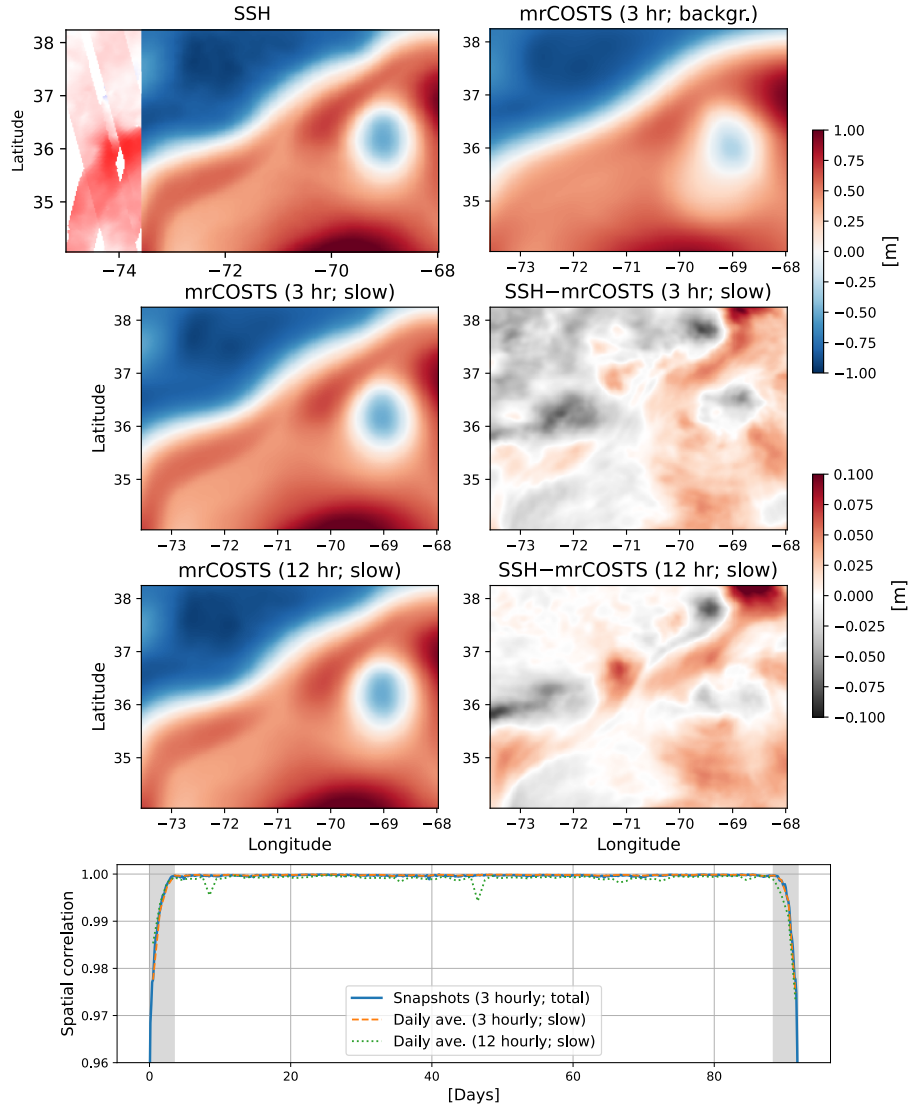


Figure 4. Instantaneous snapshots of SSHa and its mrCOSTS reconstruction on an arbitrary day. HYCOM50 output of instantaneous SSHa spatially smoothed using a Gaussian filter with a standard deviation of 10 km (a), the slowest (background) mrCOSTS band where SSHa fields were fed three hourly (b), mrCOSTS extraction of sub-inertial component (c), residual between HYCOM50 SSHa and sub-inertial component (d). The SWOT Cal/Val tracks of pass number 9 and 22 are shown in panel (a), which partially overlap with the HYCOM50 domain analyzed here. MrCOSTS extraction of the sub-inertial component where SSHa fields were fed 12 hourly (e), and its residual (f). Time series of spatial correlation between SSHa and mrCOSTS reconstructions (g). The solid blue curve documents the correlation between instantaneous SSHa and total mrCOSTS reconstruction where SSHa fields were fed three hourly. The orange-dashed and green-dotted curve shows the correlation between daily-averaged SSHa and sub-inertial mrCOSTS reconstruction where data were fed three and 12 hourly respectively. The black shading indicates the duration of COI.

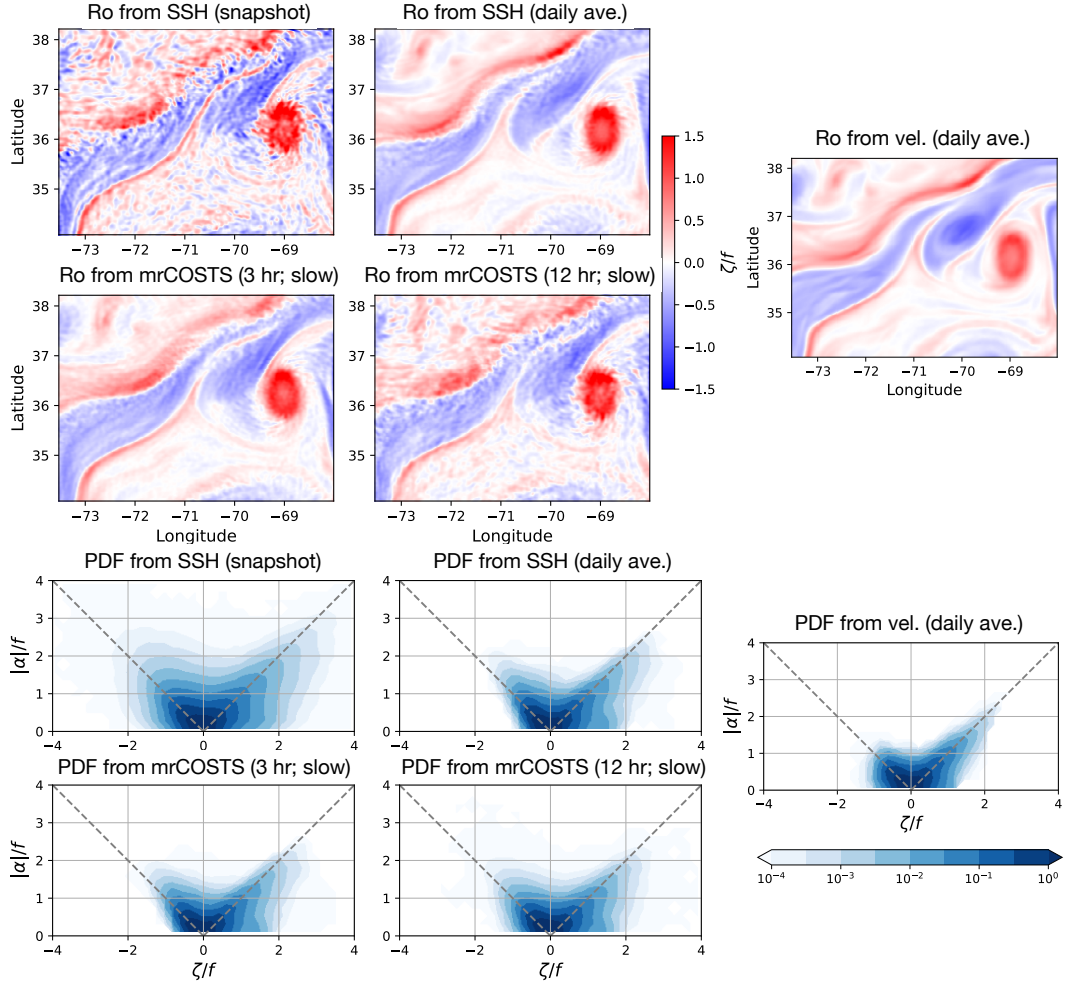


Figure 5. Spatial maps of relative vorticity normalized by the local Coriolis frequency ζ/f , viz. the local Rossby number Ro from HYCOM50. Panel (a) shows Ro diagnosed from an instantaneous SSHa field spatially smoothed using a Gaussian filter with a standard deviation of 10 km, and when the hourly SSHa fields are daily averaged to diagnose Ro (b). Instantaneous mrCOSTS reconstructions of the slow component of Ro when data are fed every three and 12 hours are documented in panels (c) and (d). Joint probability density functions (PDFs) of Ro and strain rates normalized by f for each case over the three months of August–October (e–h). A spatial map of Ro and joint PDF of Ro and strain rate at the surface computed from daily-averaged and spatially-smoothed total velocity using a Gaussian filter with a standard deviation of 10 km is shown for reference (i,j).

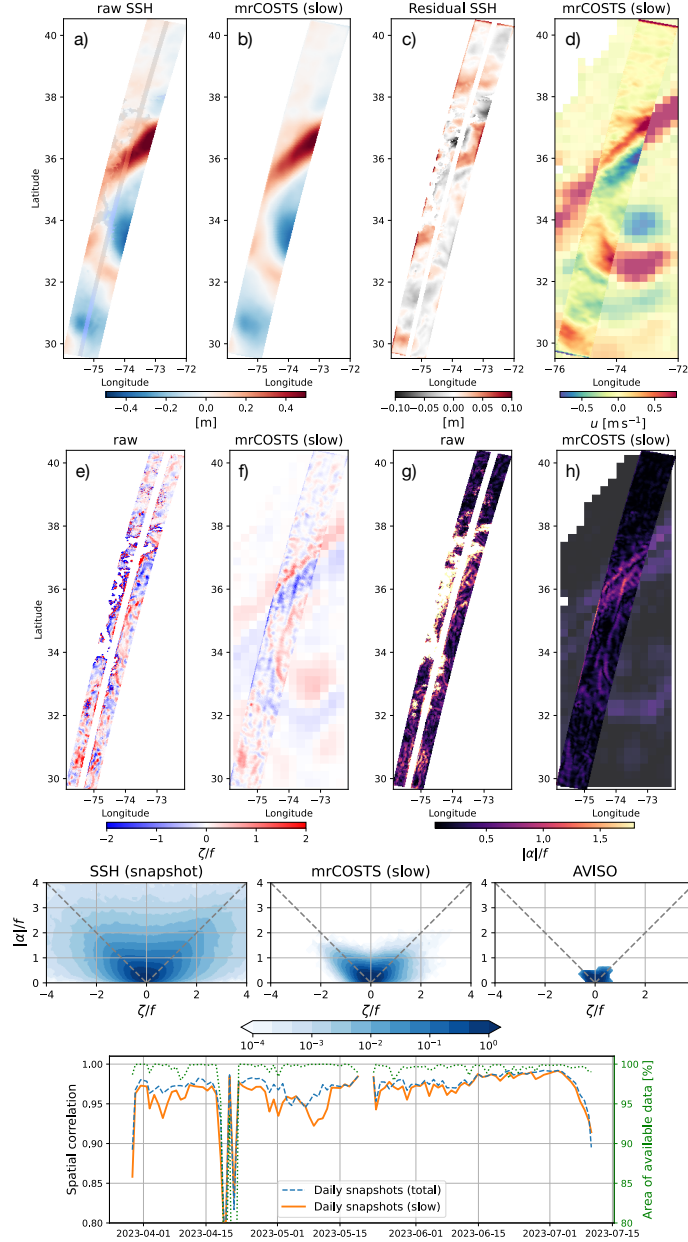


Figure 6. L3 SWOT observation of SSHa on June 21, 2023 (a), mrCOSTS reconstruction of the slow component of the spatially filtered SSHa (b), and the difference between the two (c).

Missing data and spacing between the swaths are interpolated over for SWOT SSHa and shown in a different colormap in panel (a). Zonal geostrophic velocity diagnosed from the mrCOSTS slow component (d). Relative vorticity ζ and strain rate $|\alpha|$ normalized by f diagnosed from the SWOT data and mrCOSTS slow component (e-f). Geostrophic zonal velocity, relative vorticity and strain rate from daily-averaged 0.25° gridded AVISO are shown in lighter shadings in contrast to mrCOSTS. Joint PDF of ζ/f and $|\alpha|/f$ diagnosed from raw SWOT data, mrCOSTS slow component and AVISO (i-k). The SWOT fields used in panels (a,e,g,i) were not spatially filtered as the L3 product *a priori* has some smoothing applied (Dibarboure et al., 2024).

Time series of spatial correlation between SWOT SSHa and its mrCOSTS reconstructions (l). The dashed blue curve documents the correlation between instantaneous SSHa and total mrCOSTS reconstruction. The orange-solid curve shows the correlation between SSHa and mrCOSTS slow component. The green dotted curve, plotted against the right y axis, shows the percentage of available data per snapshot.

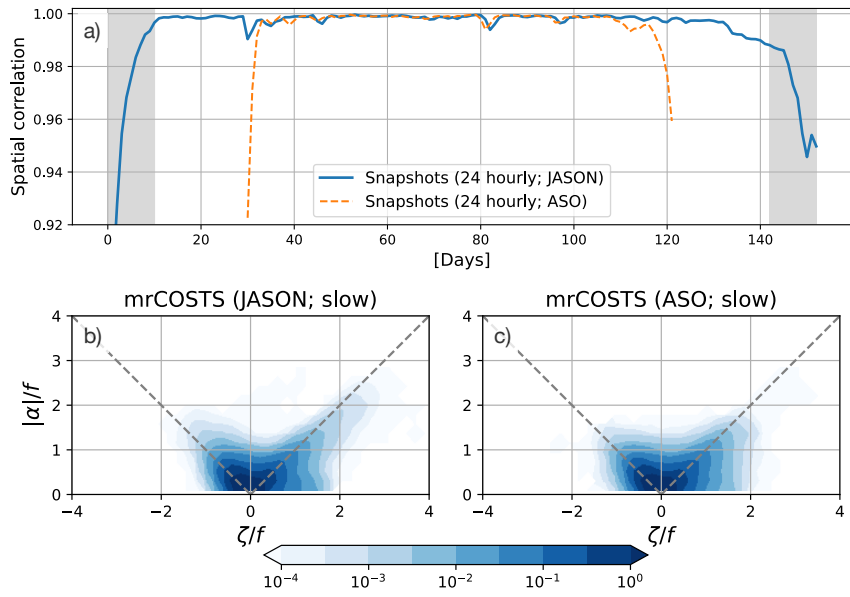


Figure 7. Time series of spatial correlation between HYCOM50 SSHa and its total reconstruction by mrCOSTS when data is fed 24 hourly (a). The black shading indicates the duration of COI for the Jul., Aug., Sept., Oct., and Nov. (JASON) case. Joint PDFs of relative vorticity and strain rate normalized by the local Coriolis frequency for the JASON (b) and ASO case (c).