

# 1 Diagnosing cross-scale kinetic energy exchanges from 2 two submesoscale permitting ocean models.

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## 7 **Key Points:**

- 8 • We used two submesoscale permitting ocean models of the North Atlantic Ocean  
9 to investigate kinetic energy exchanges at fine-scales.
- 10 • KE fluxes at fine-scales are strongly impacted by submesoscale turbulence with  
11 a stronger forward cascade in winter within the mixed-layer.
- 12 • Not accounting for ageostrophic motions yields a significant under-estimation of  
13 the forward cascade.

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**Abstract**

Fine-scale motions ( $<100$  km) contribute significantly to the exchanges and dissipation of kinetic energy in the upper ocean. However, knowledge of ocean kinetic energy at fine-scales (in terms of density and transfers) is currently limited due to the lack of sufficient observational datasets at these scales. The sea-surface height measurements of the upcoming SWOT altimeter mission should provide information on kinetic energy exchanges in the upper ocean down to 10-15 km. Numerical ocean models, able to describe ocean dynamics down to  $\sim 10$  km, have been developed in anticipation of the SWOT mission. In this study, we use two state-of-the-art, realistic, North Atlantic simulations, with horizontal resolutions  $\sim 1.5$  km, to investigate the distribution and exchanges of kinetic energy at fine-scales in the open ocean. Our results show that the distribution of kinetic energy at fine-scales approximately follows the predictions of quasi-geostrophic dynamics in summertime but is somewhat consistent with submesoscale fronts-dominated regimes in wintertime. The kinetic energy spectral fluxes are found to exhibit both inverse and forward cascade over the top 1000 m, with a maximum inverse cascade close to the average energy-containing scale. The forward cascade is confined to the ocean surface and shows a strong seasonality, both in magnitude and range of scales affected. Our analysis further indicates that high-frequency motions ( $<1$ day) play a key role in the forward cascade and that the estimates of the spectral fluxes based on geostrophic velocities fail to capture some quantitative aspects of kinetic energy exchanges across scales.

**Plain Language Summary**

The dynamics of oceanic motions with scales  $<100$  km (fine-scales) are currently not well known. This is due to the lack of sufficient observational datasets at these scales in the ocean. There are suggestions from recent studies that this class of motions impacts the distribution and exchanges of kinetic energy in the ocean. [To better understand fine-scale motions, the Surface Water and Ocean Topography \(SWOT\) satellite has been assembled. SWOT is expected for launch in 2022 and will provide an unprecedented view of the ocean down to a wavelength of 10-15 km.](#) In anticipation for SWOT mission, numerical ocean models capable of resolving fine-scales oceanic motions have been designed and implemented. In this study, we use two of these simulations to investigate how kinetic energy is exchanged between oceanic motions at fine-scales. Our results show that submesoscale turbulence (a class of oceanic turbulence at fine-scale) and high-frequency motions affects the kinetic energy exchanges by providing a route to kinetic energy towards dissipation. Also, we found that kinetic energy exchanges based on the future SWOT dataset might underestimate the true magnitude of the transfer of kinetic energy towards finer scales.

## 1 Introduction

The ocean is a turbulent fluid with a broad range of energetic scales, ranging from large  $\sim O(1000 \text{ km})$  to centimeter scales. The ocean kinetic energy is mostly concentrated in the quasi-geostrophic mesoscale eddy field with scales  $\sim O(100 \text{ km})$  (Stammer & Böning, 1992). Due to non-linear interactions among different length scales, energy can be transferred both from large to small (forward, or direct cascade) and from small to large scale (inverse cascade). Understanding the distribution of kinetic energy (KE) and variance across scales in oceanic flows is, therefore, key to our knowledge of ocean circulation (Ferrari & Wunsch, 2009).

To estimate the variance and energy associated with eddy motions at different scales, velocity wavenumber power spectral density has proven to be very efficient (Le Traon et al., 1990, 2008; Fu et al., 2010; Dufau et al., 2016; Uchida et al., 2017). However, spectral density does not indicate the direction of kinetic energy exchanges between the different scales. A better knowledge of cross-scale energy exchanges is gained by looking at the KE cascade due to nonlinearity. This important feature in turbulence study dates back to the work of Charney (1971) and Salmon (1980) on geostrophic turbulence. For stratified rotating quasi-two-dimensional fluid motion, classical geostrophic turbulence theory predicts a direct cascade of energy if the flow is depth-dependent (baroclinic) but an inverse cascade of energy if the flow is depth-independent (barotropic). In particular, for the ocean with a surface intensified stratification, energy from higher baroclinic modes concentrates in the first mode and then converges toward the scale of the Rossby radius of deformation ( $R_d$ ) (Smith & Vallis, 2002). At  $R_d$ , baroclinic energy is converted to barotropic mode via barotropization. At this point, most of the energy near the deformation scale cascade towards larger scales while a small fraction undergoes direct cascades to dissipation (see Figure 1).

This prediction of geostrophic turbulence theory has been observed both in numerical simulations and the real ocean but with a little discrepancy. Based on altimeter data, Scott and Wang (2005) showed that an inverse cascade of energy dominates the (Pacific) ocean at scales larger than  $R_d$ . So, if one agrees that the altimeter data is reflecting the first baroclinic mode (Smith & Vallis, 2002), then this is in contrast with geostrophic turbulence theory which predicts a forward cascade for a baroclinic flow. Scott and Wang (2005) argued that there might be an inverse cascade associated with the first baroclinic mode and that this would only partially reduce the forward flux of total baroclinic energy. This total energy forward flux is the source of the kinetic energy that arrives near the deformation scale from the large-scale mean flow via baroclinic instability. From this discrepancy, two questions arise. (i) is the inverse cascade seen at the surface due to the barotropic mode? or (ii) is it possible that the baroclinic modes experience an inverse cascade? Scott and

85 Arbic (2007) using a 2-layer model simulation showed that the inverse cascade at the ocean  
 86 surface is mostly baroclinic with a small contribution from the barotropic mode. The results  
 87 from Scott and Arbic (2007) are consistent with the proposed modification to geostrophic  
 88 turbulence by Scott and Wang (2005). More recent literature (Schlösser & Eden, 2007;  
 89 Sasaki et al., 2017; Tulloch et al., 2011; Aluie et al., 2017; Brüggemann & Eden, 2015;  
 90 Kjellsson & Zanna, 2017; Khatri et al., 2018) have also shown that an inverse cascade of  
 91 energy mostly dominates the surface ocean at scales larger than  $R_d$ .

92 In contrast, little is known regarding energy cascade at scales  $< R_d$ , where oceanic  
 93 motion is dominated by submesoscale motions ( $< 20 - 50$  km). Results from numerical sim-  
 94 ulation and observation have shown an injection of energy in winter at submesoscale (Sasaki  
 95 et al., 2017). This energy injection is partly responsible for both meso and submesoscale  
 96 seasonality (Uchida et al., 2017; Capet, Campos, & Paiva, 2008; Sasaki et al., 2014) and has  
 97 been argued to be associated with mixed layer instability (Callies, Ferrari, et al., 2015; Qiu  
 98 et al., 2014; Sasaki et al., 2014; Brannigan et al., 2015; Rocha et al., 2016). This seasonality  
 99 is responsible for the shallowing of KE spectral slope from -3 in summer to -2 in winter and  
 100 is usually interpreted as a shift from turbulence dominated by interior gradients (Philips  
 101 regime) to a regime dominated by surface driven turbulence (Charney regime) (Sasaki et  
 102 al., 2014). Apart from the work of Sasaki et al. (2014, 2017); Schubert et al. (2020), we are  
 103 unaware of any investigation on the implication of submesoscale seasonality on cross-scale  
 104 energy exchanges at the basin scale and one of the objective of this study is to investigate  
 105 how submesoscales modify cross-scale kinetic energy exchanges at fine-scales.

106 Submesoscale resolving ocean models have been developed in anticipation of the Surface  
 107 Water and Ocean Topography (SWOT) satellite mission (Fu et al., 2010). On a global scale,  
 108 satellite altimeters remain the primary source of information on the distribution of energy  
 109 across scales. However, at the moment, the resolution capability of our existing ocean-  
 110 observing satellite altimeters stands at roughly 70 km (Dufau et al., 2016). This limitation  
 111 undermines our ability to investigate energy exchanges at scales  $< 100$  km. To solve this  
 112 challenge, SWOT is implemented to provide ten times higher resolution than conventional  
 113 altimeters, and numerical ocean models have been designed to prepare for SWOT. These  
 114 state-of-the-art numerical experiments with high-resolution capability, thereby provide an  
 115 opportunity to study cross-scale energy exchanges down to kilometeric scales.

116 In this study, we aim to investigate the distribution and transfer of energy across  
 117 different scales by using outputs of two submesoscale permitting ocean models of the North  
 118 Atlantic. In particular, we focus on the seasonality and depth penetration of cross-scale  
 119 KE variance and transfer with an emphasis on scales  $< 100$  km. This paper is organized  
 120 as follows; section 2 presents a description of the two numerical simulations. In section 3,

121 we examine the kinetic energy wavenumber spectral density and slope. The KE cascade,  
122 its seasonality, and the role of high frequency and ageostrophic motions on the cascade are  
123 discussed in section 4. Finally, in section 5, we summarize the findings and discuss the  
124 relevance of this work in anticipation for SWOT mission.

## 125 **2 Numerical simulations of the North Atlantic Ocean**

126 In this study, we use numerical outputs from two submesoscale eddy-permitting sim-  
127 ulations of the North Atlantic: a NEMO-based simulation with a horizontal resolution of  
128  $1/60^\circ$  (NATL60) and an HYCOM-based (HYbrid Coordinate Ocean Model) simulation with  
129 a horizontal resolution of  $1/50^\circ$  (HYCOM50).

130 The NEMO-based NATL60 has a horizontal grid spacing ranging from 1.6 km at  $26^\circ\text{N}$   
131 to 0.9 km at  $65^\circ\text{N}$ . The initial and open boundary conditions are based on the GLORYS2v3  
132 ocean reanalysis with a relaxation zone at the northern boundary for sea-ice concentration  
133 and thickness. The model has 300 vertical levels with a resolution of 1 m at the top-most  
134 layers. The grid and bathymetry follow Ducouso et al. (2017), while the atmospheric forcing  
135 is based on DFS5.2 (Dussin et al., 2018). DFS5.2 forcing is based on ERA-interim reanalysis.  
136 The spatial resolution of the atmospheric fields is 0.75 degrees. All variables used to compute  
137 turbulent fluxes (air temperature and humidity at 2m, wind velocity components at 10m)  
138 are 3-hourly. In order to implicitly adapt lateral viscosity and diffusivity to flow properties,  
139 a third-order upwind advection scheme is used for both momentum and tracers in the model  
140 simulation. The model was spun-up for six months, and a one-year simulation output from  
141 October 2012 to September 2013 is used in this study. The simulation output used in this  
142 study is the same as the one used in Amores et al. (2018), Buckingham et al. (2019) and  
143 Ajayi et al. (2020). An earlier version of this simulation set-up was used in Ducouso et al.  
144 (2017) and Fresnay et al. (2018).

145 NATL60 ocean simulation has been evaluated, using in situ observations in terms of the  
146 kinetic energy levels at different wavelengths (see Figure 1 of supporting information docu-  
147 ment (SI)). In terms of the dynamics of the resolved fine-scales in the upper ocean, results  
148 from Buckingham et al. (2019) show that the statistics of horizontal velocity tensor predicted  
149 by NATL60 agree reasonably well with observation (OSMOSIS datasets). However, their  
150 results also show that there is a likelihood of extreme divergent motions in OSMOSIS that  
151 is not captured by NATL60. That NATL60 underestimates divergent motions compared to  
152 observation isn't that surprising because NATL60 model simulation is without tidal forcing,  
153 one of the major sources of wave energy. The model, however, reproduces fairly well other  
154 forms of internal gravity waves (see Figure 2 of SI).

155 The HYCOM-based HYCOM50 extends from 28°S to 80°N and has a horizontal grid  
 156 spacing ranging from 2.25 km at the equator,  $\sim 1.5$  km in the Gulf Stream region, and 1  
 157 km in the subpolar gyre. As for NATL60, the effective resolution is about 10–15 km. The  
 158 vertical coordinate is hybrid and consists of 32 layers. The simulation is initialized using  
 159 potential temperature and salinity from the GDEM climatology and spun up from rest for  
 160 20 years using climatological atmospheric forcing from ERA-40 (Uppala et al., 2005), with  
 161 3-hourly wind anomalies from the Fleet Numerical Meteorology and Oceanography Center  
 162 3-hourly Navy Operational Global Atmospheric Prediction System (NOGAPS) for the year  
 163 2003. The horizontal viscosity operator is a combination of Laplacian and Biharmonic. The  
 164 bathymetry is based on the Naval Research Laboratory (NRL) digital bathymetry database.  
 165 The model configuration and a detailed evaluation of the model results in the Gulf Stream  
 166 region with observations are documented in Chassignet and Xu (2017).

167 In this study, we use the output from October 2012 to September 2013 for NATL60  
 168 and year 20 (last year of the simulation) for HYCOM50. Winter and summer correspond  
 169 to January-February-March (JFM) and July-August-September (JAS), respectively. The  
 170 summer analysis presented for NATL60 corresponds to that of the year 2013. Since NATL60  
 171 covers a smaller domain than HYCOM50, we consider the HYCOM50 outputs for the same  
 172 region covered by NATL60 to have comparable results. To capture regional variability in  
 173 the distribution of energy across scales, we perform spectral analysis in sub-domains of  
 174  $14\ 10^\circ \times 10^\circ$  boxes across the North Atlantic. We focus specifically on quantifying kinetic  
 175 energy wavenumber spectral density (Eq. 1) and flux (Eq. 2) using horizontal velocity  
 176 fields. In equation (1) and equation (2),  $\hat{\cdot}$  refers to Fourier transform,  $*$  represents the  
 177 complex conjugate,  $Re$  refers to the real part of a complex number and  $k = \sqrt{k_x^2 + k_y^2}$ .  
 178 Before performing spectral analysis, the 2D velocity field from each subdomain (box) is  
 179 detrended in both directions, and a 50% cosine taper window (turkey windowing) is applied  
 180 for tapering. An FFT is applied to the tapered data, and a 1D isotropic spectrum is obtained  
 181 by averaging in the azimuthal direction. Our spectral method is consistent with procedures  
 182 previously used in Stammer and Böning (1992); Sasaki and Klein (2012); and Chassignet  
 183 and Xu (2017).

$$E(k) = \int_k^{k+\delta k} [\hat{\mathbf{u}}^* \cdot \hat{\mathbf{u}}](k) dk \quad (1)$$

$$\Pi(k) = \int_k^{k_s} -Re \left[ \hat{\mathbf{u}}^* \cdot \left( \mathbf{u} \cdot \widehat{\nabla_{\mathbf{H}} \mathbf{u}} \right) \right] (k) dk \quad (2)$$

184 Both NATL60 and HYCOM50 resolve the first Rossby radius of deformation everywhere  
 185 within the model domain, and these simulations reproduce realistic eddy statistics with levels

186 of kinetic energy in the range of altimetric observations (Chassignet and Xu (2017), see also  
 187 Figure 2). A summary of the model parameters is tabulated in Table 1. Both simulations  
 188 are submesoscale permitting ocean models and a discussion on the ability of these models  
 189 to resolve the dynamics of submesoscales in terms of eddy length scale, submesoscale energy  
 190 and their associated seasonality can be found in Ajayi et al. (2020).

191 We present in Figure 3, the root-mean-square of sea surface height computed from  
 192 one year daily outputs of NATL60, HYCOM50, compared with AVISO. The AVISO SSH  
 193 field is derived from observations obtained by altimeter missions and then interpolated  
 194 onto a  $0.25^\circ$  mercator grid. In this comparison, we have used AVISO mean dynamical  
 195 topography dataset from October 2012 to September 2013. The SSH fields in Figure 3  
 196 from NATL60 and HYCOM50 have been resampled onto a  $0.25^\circ$  grid to have comparable  
 197 results. The SSH rms of the models and AVISO data appears to be consistent in terms  
 198 of the geographical pattern of energetic oceanic motions except for the differences along  
 199 the Gulf Stream. The Gulf Stream seems to be more energetic in HYCOM50 compared to  
 200 NATL60 and AVISO. HYCOM50 has a spin-up of 20 years while NATL60 has a spin-up  
 201 of 6 months. We hypothesise that HYCOM50 long spin-up allows for the full development  
 202 of the Gulf Stream energetics and this difference in terms of spin-up could be contributing  
 203 to differences in the overall energetics of the two simulations. The implication of the short  
 204 spin-up for NATL60 is more obvious in the time evolution of the domain-averaged kinetic  
 205 energy (see Figure 3 of SI). The kinetic energy is still increasing with time and is yet to  
 206 attain equilibrium. A similar curve for HYCOM50 is available in Figure 2 of Chassignet  
 207 and Xu (2017). The comparison of the KE spectral in summer of the year 2012 vs the year  
 208 2013<sup>1</sup> (see Figure 4 of SI) further highlights the increase in energy (for NATL60) that is  
 209 characterized by higher variance and larger eddies at low wavenumbers in year 2013.

210 The two simulations are similar but are not without differences. HYCOM50 appears  
 211 to be more energetics compared to NATL60. The disparity between the two models' energy  
 212 level is not the main focus of this paper, but we shall propose a few reasons why the two  
 213 models could differ in terms of energetics. Firstly, we hypothesise that the eddy structures  
 214 in NATL60 are not fully developed due to the short spin-up (6 months for NATL60 versus  
 215 20 years for HYCOM50). The first two years of the HYCOM50 simulation show an increase  
 216 of total kinetic energy level; see Figure 2 in Chassignet and Xu (2017). Furthermore,  
 217 the typical scale of eddies are smaller in NATL60 compared to HYCOM50 (Ajayi et al.,  
 218 2020), and this could be a direct consequence of the shorter duration of the model spin-up.  
 219 Secondly, the question did arise as to whether the coarser vertical resolution in HYCOM50

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<sup>1</sup> The summer analysis presented in the article corresponds to that of year 2013

220 (32 hybrid vertical layers versus 300 z-levels in NATL60) could lead to a stronger inverse  
 221 cascade and hence a higher energy level because of an under-resolved stratification and the  
 222 depth dependence of flows. A comparison of the vorticity spectral coherence as a function  
 223 of depth shows that the two simulations are essentially identical in terms of the depth  
 224 penetration of energetic eddy structures (Ajayi et al., 2020). Furthermore, in section 4, a  
 225 comparison of the KE spectral flux at depths for the two simulations will show that the  
 226 HYCOM50 upscale energy flux is not surface intensified and that having only 32 isopycnal  
 227 vertical levels is not detrimental to the representation of the dynamics in the ocean interior.  
 228 Thirdly, the choice of sub-grid parameterization is different between the two simulations  
 229 and could have a substantial effect on how energy is dissipated in each model.

230 Up until recently, most basin-scale numerical models like those used in this study usu-  
 231 ally store simulation outputs in the form of daily averages due to limitations in storage  
 232 and computational resources. This limitation comes with a caveat. Daily averaging the  
 233 model outputs suppresses high-frequency motions ( $> f$ , where  $f$  is the Coriolis frequency).  
 234 These motions are mostly dominated by ageostrophic motions that include unbalanced sub-  
 235 mesoscales and fast propagating internal gravity waves. There are new shreds of evidence,  
 236 based on idealized simulations that suggest that high-frequency motions (particularly wave  
 237 motions) can provide a route to kinetic energy dissipation (Barkan et al., 2017; Rocha et  
 238 al., 2018). It would be interesting to investigate these new results in realistic simulations  
 239 like NATL60 and HYCOM50. Unfortunately, NATL60 and HYCOM50 have most of their  
 240 outputs stored in daily averages except for the last month of simulation for HYCOM50 and  
 241 surface quantities for NATL60. To that end, the results presented in this study are based  
 242 on daily averages of velocity fields from the two simulations. We have only used analysis  
 243 based on hourly output where necessary to illustrate the impacts of high-frequency motions  
 244 on the kinetic energy distribution and exchanges.

### 245 **3 Distribution of Kinetic Energy**

246 In this section, we discuss the variance at different scales of motions by analysing  
 247 the kinetic energy wavenumber power spectral density. In general, horizontal wavenumber  
 248 spectral density exhibits power-law behavior, where the exponent is interpreted in terms of  
 249 the dynamical processes governing the eddy energy transfer. Existing theoretical frameworks  
 250 (for horizontal velocity at scale  $> Rd$ ) predict a spectral slope of -3 and -5/3 for QG and  
 251 SQG (surface quasi-geostrophic) turbulence respectively. A slope of -2 is also well known for  
 252 a front dominated flow (Shcherbina et al., 2013; Callies & Ferrari, 2013). Over the years,  
 253 many research works have tried to establish the accuracy of these predictions by using  
 254 outputs of realistic ocean models (Sasaki & Klein, 2012; Chassignet & Xu, 2017; Uchida et



al., 2017) and also recently within the context of the real ocean by using altimeter dataset (Le Traon et al., 1990; Dufau et al., 2016). Their results have argued for the non-existence of a universal wavenumber spectrum (Le Traon et al., 2008) following observed regional variability. We shall discuss in section 3.1, the distribution of kinetic energy as predicted by NATL60 HYCOM50 and in section 3.2 we discuss the variability associated with this distribution and their associated estimated slope.

### 3.1 Spectral Density

In Figure 4, we present the kinetic energy spectral density as a function of depth for the two simulations. For simplicity, we show this comparison only for Box 8, a box located at the center of the North Atlantic basin. In most of the regions, the peak of the spectral density is around the mesoscale motions (100–500 km). As expected, the energy associated with large-scale motion is relatively higher than that of fine-scales. The peak of the spectral density is preserved with depth, while the variance at all scales decreases with depth. The comparison between the two simulations is illustrated better in Figure 5a where we present the depth-averages of annual KE spectral density for the two simulations in the same region (Box 8). The spectral densities from the two models agree well with an approximate slope of  $-3$ , a value that is characteristic for Quasi-Geostrophic (QG) prediction. In this QG regime, submesoscale structures are expected to be weakly energetic while energy is concentrated at the mesoscales. The depth-averages of the winter and summer KE spectral density are presented in Figure 5b. There is a seasonality in the spectral density that is mostly associated with an increase in the variance at submesoscales in winter. The spectral shape in both winter and summer are somewhat QG. We find this surprising given that previous studies (Callies & Ferrari, 2013; Shcherbina et al., 2013; Sasaki et al., 2014) have shown that in the presence of intense submesoscales in winter, KE spectral density is likely to have  $k^{-2}$  or  $k^{-5/3}$  spectral shape.

As previously highlighted, daily averaging the velocity fields before spectral estimation could suppress the signature of high-frequency motions (unbalanced submesoscales, ageostrophic wave motions) and this could affect the distribution of energy implied from this sort of spectral analysis. To ascertain this, we compared NATL60 surface KE spectral density for hourly averages versus daily averages in Figure 6a. This comparison is only possible for NATL60 because HYCOM50 dataset is stored in daily averages except for the last month of the simulation. In this comparison for NATL60 (Figure 6a), submesoscale motions are more energetic than presented in the daily spectral density. There is a tendency for the distribution (annual average) of energy at submesoscales in the hourly spectral to follow a  $k^{-2}$  spectra shape, a characteristics of a regime associated with fronts.

290 The winter versus summer spectral density computed from NATL60 hourly averages  
 291 of the velocity field (Figure 6b), clearly show how high-frequency motions and energetic  
 292 submesoscales drive the seasonality of kinetic energy distribution at fine-scales. There is a  
 293 shift in the spectral slope from  $-3$  in summer to  $-2$  in winter. This shift can be interpreted  
 294 as a change in dynamics from the interior QG Phillips-like regime to a surface intensified  
 295 Charney-like regime (Sasaki et al., 2014). In this Charney-like regime, submesoscales are  
 296 associated with large vertical velocity and, in turn, large submesoscale buoyancy fluxes that  
 297 are suggested to feeds mesoscales through an inverse kinetic energy cascade. This result is  
 298 consistent with the findings of Sasaki et al. (2014); Callies and Ferrari (2013), and it further  
 299 evaluates the ability of NATL60 to resolves physical processes at fine-scales. In section (4),  
 300 we shall also discuss how this regime change from summer to winter affects the redistribution  
 301 of kinetic energy.

### 302 3.2 Spectral Slope

303 As we have shown in the previous section, a quick way to estimate the wavenumber  
 304 spectral power law is to compute the 1D wavenumber spectral density then estimate a slope  
 305 from this spectral by fitting a line to the spectral density curve within a selected wavenumber  
 306 range. This method is fast and easy to implement and provides a way to investigate regional  
 307 variability of ocean energetics both at the basin and global scale. For studies on mesoscale  
 308 energetics using satellite datasets and model outputs, this wavenumber range is mostly  
 309 within the error limits of the altimeter instrument ( $\sim 70$  km) and the horizontal scale of  
 310 meso/large scale motion (250–300 km). One drawback of this approach is that it does not  
 311 account for the changes in the scale of average energetic eddy structures with latitude. Scales  
 312 of motions that are mesoscales in the polar regions could be classified as submesoscales in  
 313 the tropics.

314 To characterize the spectral signature correctly, several recent studies have tried to  
 315 propose different approaches to estimate the wavenumber spectral power law. For mesoscale  
 316 resolving altimetry datasets, Vergara et al. (2019) estimated spectral slope between the peak  
 317 of the spectral and the minimum of the Rossby radius and the Rhines scale following Eden  
 318 (2007). A similar approach was presented in Sasaki and Klein (2012), where the authors  
 319 estimated spectral slope between a fixed wavelength of 30 km (at the lower bound), and a  
 320 scale that corresponds to the peak of the KE wavenumber spectral.

321 In order to show how sensitive the estimated slopes are to the selected wavelength range,  
 322 we present in Figure 7a the average KE wavenumber spectral density and slope for box 3  
 323 in March for three different selected wavenumber ranges. The dashed lines with colors  
 324 red, blue, and black represent the 10–100 km, 10–250 km, and 70–250 km, wavelength,

325 respectively. The estimated slopes for these three different wavelength ranges have different  
 326 values, therefore raising the question as to which slope is most representative of the dynamics  
 327 of this region. We repeat this analysis for all the boxes and present the map in Figure 8. The  
 328 mismatch is particularly pronounced in the sub-polar region, where the scales of the eddy  
 329 structures are relatively smaller. The 70–250 km wavelength range is a typical wavelength  
 330 for estimating spectral slope for satellite datasets because 70 km roughly corresponds to  
 331 the wavelength where the satellite data becomes noisy. The spectral slope in this range is  
 332 consistent with the already published work of Dufau et al. (2016) and Chassignet and Xu  
 333 (2017).

334 To avoid the sensitivity of the estimated spectral slope to an a-priori selected wavelength  
 335 range, we introduce an approach that takes into account the dynamics of the regions and  
 336 the resolving capability of the model by estimating the spectral slope (Figure 7b) between  
 337 the energy-containing scale (Kjellsson & Zanna, 2017) and the effective resolution of the  
 338 model (Soufflet et al., 2016). The energy-containing scale (which represents the scale of the  
 339 most energetic eddy structure) is estimated from the kinetic energy wavenumber spectral  
 340 using equation (3) while the effective resolution (a function of the model grid-size) is taken  
 341 as  $5 \times$  the model grid size, which is roughly equal to 10 km for both models. This approach  
 342 takes into account the scale of the energetic eddy structures within the flow region and also  
 343 takes into account the geographical variability of this scale, and therefore provides a way to  
 344 infer dynamical properties of oceanic motions in different regions.

$$\lambda_e = \frac{\int \int E(k_x, k_y) dk_x dk_y}{\int \int \sqrt{k_x^2 + k_y^2} E(k_x, k_y) dk_x dk_y} \quad (3)$$

345 We apply this technique to the output of both simulations (KE spectral density from  
 346 daily averages), and we present the estimated spectral slope and the energy-containing scale  
 347 (integral scale) for all the boxes in the North Atlantic (Figure 9). The estimated integral  
 348 scale from the wavenumber spectral density represents the averaged scale of energetic struc-  
 349 tures in the selected region. On one hand, this scale varies regionally and fairly follows the  
 350 variability of the Rossby radius of deformation with latitude, with high values in the south  
 351 and relatively low values in the north. On the other hand, the estimated slope across the  
 352 basin is almost uniform and follows the prediction of QG with a slope value  $\sim k^{-3}$ . This  
 353 consistency with the QG prediction is observed in both model outputs and also holds in the  
 354 well known high energetic Gulf stream (box 1) and the low energetic OSMOSIS (box 10)  
 355 regions. This result from daily fields comes with the caveat of suppressing the impact of  
 356 high-frequency motions on the estimated slope values.

357 In order to investigate whether accounting for high-frequency motions would affect  
 358 our estimation of spectral slopes, we estimate the spectral slope (from KE spectral density  
 359 computed using hourly averages of velocity fields). We do this only for NATL60 because  
 360 HYCOM50 surface hourly output is available only for one month. The NATL60 spectral  
 361 slope from hourly spectral density in all the regions (Figure 10) has smaller values than  
 362 that of daily averages. As previously highlighted in section 3.1, this implies that the actual  
 363 spectral are shallower in hourly fields (as a result of stronger variance at the fine-scales)  
 364 compared to daily fields. The impact of intense submesoscale and ageostrophic flows is  
 365 better illustrated in the winter/summer map of the spectral slope (Figure 11). The slopes  
 366 in winter have values that are closer to  $k^{-2}$  throughout the domain. This indicates that  
 367 high-frequency motions are quite significant in the North Atlantic ocean basin. In the next  
 368 section, we shall discuss the impact of different dynamics on the redistribution of kinetic  
 369 energy.

## 370 4 Kinetic Energy Cascade

371 In this section, we present and discuss the exchange of energy due to non-linearity  
 372 across different scales of motion. This exchange is estimated from the horizontal velocity  
 373 fields using equation (2). A positive flux represents a direct cascade of energy, while a  
 374 negative value represents an inverse cascade of energy.

### 375 4.1 Annual Averages of Kinetic Energy Cascade

376 We show in Figure 12 the KE spectral flux computed using one year's daily outputs  
 377 of surface velocity fields. For simplicity, we show plots for boxes 3, 8, and 11 representing  
 378 latitudes of 35°N, 45°N, and 55°N, respectively. In all the boxes and both models, the  
 379 spectral flux is dominated by an inverse cascade of energy at large scales (between 25-50  
 380 km and 500 km) and a forward cascade of energy below 25-50 km. As observed earlier,  
 381 the spectral slope from daily averaged fields has a value that is  $\sim -3$ , a characteristic  
 382 of QG turbulence. The energy exchanges computed using the same data show that the  
 383 flux is mostly upscale with a little forward flux at fine-scales. **While the inverse cascade  
 384 is a well known phenomenon in geostrophic turbulence, the dynamics responsible for the  
 385 forward cascade at fine-scales is, however, still a subject for discussion.** One would expect a  
 386 forward flux at fine-scales to be accompanied by a ( $k^{-2}$  or  $k^{-5/3}$ ) spectral slope. This sort of  
 387 relationship between spectral density and flux is not observed for estimates coming from the  
 388 daily fields. This finding is, however not new and agrees with the results of Brüggemann and  
 389 Eden (2015). The authors show that as soon as ageostrophic dynamics become important  
 390 in a quasi-QG flow, the flow is no longer restricted to an inverse cascade of energy. Instead,

391 the kinetic energy can proceed toward smaller scales providing a direct route to dissipation.  
 392 We believe the fine-scale ageostrophic motions resolved by these simulations are significant  
 393 for the flow to be in a regime that supports a forward cascade of energy.

394 It is interesting to note that the scale at which the inverse cascade is most intense  
 395 coincides with the energy-containing scale (dashed line in Figure 12 estimated from the  
 396 kinetic energy spectral density). This signifies that the inverse cascade is maximum at the  
 397 scale of the most energetic eddies. Just like the energy-containing scale, the scale of the most  
 398 intense inverse cascade varies with latitude with relatively smaller values in the sub-polar  
 399 regions (Figure 13). Also, depending on the region, part of the submesoscale range (0 to 25-  
 400 50 km) falls to the left of the zero-crossing (where the flux changes sign). This implies that  
 401 submesoscales motions are involved in fluxing energy to large-scale via an inverse cascade  
 402 of energy. This scale of zero-crossing varies across the basin with higher values in high  
 403 EKE regions and relatively smaller values in low EKE regions (Figure 13). This regional  
 404 variability in the value of the zero-crossing shows that at fine-scales, the spatial scale at  
 405 which energy is injected into the flow is greater in eddy-rich regions, compared to eddy-poor  
 406 regions.

407 Figure 14 presents the KE spectral flux for Box 8 as a function of depth. The overall  
 408 shape of the flux is preserved, and the scale at which the inverse cascade is maximum is  
 409 also consistent with depth. The strength of the inverse cascade decreases with depth, and  
 410 the direct cascade at high wavenumbers is confined mostly to the surface. In section 3.1,  
 411 the question was raised as to whether HYCOM50's higher KE, when compared to NATL60,  
 412 is a consequence of HYCOM50's coarser vertical resolution, which could lead to a surface  
 413 intensified inverse cascade and hence more energetic surface eddies. However, in Figure 14,  
 414 we can see that both at the surface and all depth levels, the estimated inverse and direct  
 415 cascade is stronger in HYCOM50 than in NATL60. This indicates that HYCOM50 upscale  
 416 energy flux is not surfaced intensified and we can conclude that the disparity between the  
 417 two models in terms of energy levels is most likely due to differences in the length of the  
 418 spin-up [as well as the choice of sub-grid scale parametrization](#), and that having only 32  
 419 isopycnal vertical levels is not detrimental to the representation of the dynamics in the  
 420 ocean interior.

## 421 **4.2 Seasonality of Energy Cascade**

422 In this section, we present the seasonality of the kinetic energy spectral flux by compar-  
 423 ing winter (JFM) and summer (JAS) averages. Figure 15 shows the winter cascade (in blue)  
 424 and the summer cascade (in red). There are two notable differences between the seasons.  
 425 First, there is a shift in the zero crossings to higher wavenumbers in winter. Second, there

426 is a stronger forward cascade within the submesoscale range in winter. As highlighted in  
 427 the preceding section, a zero-crossing at the high wavenumbers partly indicates how much  
 428 submesoscale motions are involved in feeding large-scale motions via an inverse cascade of  
 429 energy. So, a shift to higher wavenumbers in wintertime signifies that smaller-scale struc-  
 430 tures are involved in fluxing energy to larger scales. It is noteworthy that the integral scale  
 431 and scale of the maximum inverse cascade also undergo seasonality. There is a shift in the  
 432 scale to high wavenumber from winter to summer. This can be interpreted as a reduction  
 433 in the average size of energetic eddies structures in winter. This sort of seasonality in eddy  
 434 length-scale is documented in Ajayi et al. (2020).

435 It is interesting to understand the contribution of the different dynamics on the sea-  
 436 sonality of the energy cascade. Recent studies have shown that submesoscales are energetic  
 437 in wintertime (Mensa et al., 2013; Sasaki et al., 2014; Callies, Flierl, et al., 2015; Rocha  
 438 et al., 2016), and their emergence is forced by mechanisms such as frontogenesis, wind-  
 439 induced frontal instabilities, mixed layer instability among many others (Thomas, 2008;  
 440 McWilliams, 2016). Sasaki et al. (2014) argued that submesoscales generated via mixed  
 441 layer instability could feed large scale motion via an inverse cascade of energy, hence a shift  
 442 in the zero-crossing towards high wavenumbers. More recently, Schubert et al. (2020) used  
 443 both coarse-graining approach (Aluie et al., 2017) and spectral analysis to investigate the  
 444 role of mixed layer baroclinic instabilities on kinetic energy exchanges. Their results show  
 445 that mesoscale oceanic eddies are strengthened by the absorption of submesoscale mixed  
 446 layer eddies and that the forward cascade of energy at very fine-scales occur mostly in fron-  
 447 togenic regions. Following these aforementioned findings, we hypothesize that the increased  
 448 forward cascade presented in this study, could be associated with frontogenesis and subme-  
 449 soscale frontal instabilities. This seasonality highlights how submesoscale motions modulate  
 450 the redistribution of energy between scales of motions; hence, the need for climate (ocean)  
 451 models with submesoscale resolving capability.

### 452 **4.3 Impact of High-Frequency Motions on Energy Cascade**

453 In section 3.1, we observed that in the presence of high-frequency motions, the spectral  
 454 densities (computed from hourly fields) are shallower with an increased variance at fine-  
 455 scales compared to daily averaged fields. In this section, we are going to discuss the impact  
 456 of this increased variance on energy exchanges.

457 The contribution of high-frequency motions to kinetic energy spectral flux is highlighted  
 458 in Figure 16a where we show the comparison of the spectral flux computed from daily  
 459 versus hourly fields. The magnitude of the forward cascade at submesoscales is significantly  
 460 stronger in hourly spectral flux. That high-frequency motions can provide a pathway to

461 kinetic energy dissipation is illustrated in these results. The dynamics responsible for this  
 462 increase in forward cascade are likely due to energy sink by generated internal gravity waves  
 463 and unbalanced submesoscales. It is interesting to note that the scale of the most intense  
 464 inverse cascade remains the same while there is a slight shift in the zero-crossing towards  
 465 higher wavenumber. This result suggests that the impact of high-frequency motion on  
 466 energy exchanges is mostly concentrated at fine-scales.

467 As seen in the daily averages, the winter flux is equally stronger in hourly fields (Figure  
 468 16a). We believe that this increase is attributed to resolved internal gravity waves and  
 469 intense (un)balanced submesoscales. However in summer time both daily and hourly flux  
 470 are identical.

#### 471 4.4 Diagnosing Spectral Flux from SWOT

472 NATL60 and HYCOM50 are submesoscale permitting model simulations that have  
 473 been created to simulate the scales of motions that we expect SWOT to see from space.  
 474 SWOT will provide measurements of sea surface heights from which velocities (based on  
 475 geostrophic approximations) will be inferred. Geostrophically balanced motions dominate  
 476 the ocean at meso/large-scale, and the inferred geostrophic velocities at this scale mostly  
 477 reflect the absolute velocity of these large scales motions. However, geostrophy is less accu-  
 478 rate for fine-scale motions, particularly at the submesoscales and this remains a challenge  
 479 due to the projection of ageostrophic motions on SSH field. [Recent studies have shown](#)  
 480 [that some classes of non-wave ageostrophic motions could impact the forward cascade of](#)  
 481 [kinetic energy at fine-scales \(Capet, McWilliams, et al., 2008\)](#). We have equally shown in  
 482 the previous sections that high-frequency motions in the form of ageostrophic waves motion  
 483 and unbalanced submesoscales can contribute significantly to a forward cascade of energy at  
 484 fine-scales. In light of this, we are not sure (if by using satellite datasets), we can accurately  
 485 estimate the redistribution of kinetic energy at fine-scales. Given that SWOT will provide  
 486 information down to  $\sim 15$  km. We are curious to see if the geostrophically inferred surface  
 487 velocity would capture the accurate energetics at scales  $< 50$  km where geostrophy is likely  
 488 to fail.

489 To investigate this, we present in Figure 17 the spectral flux from total velocity and  
 490 geostrophic velocity for three regions (same as for the previous sections). The geostrophic  
 491 velocity is estimated from the sea surface height (SSH) using the geostrophic approximation.  
 492 The strength of the energy cascade differs between the flux computed from the total velocity  
 493 and that of the geostrophic velocity. This difference is consistent in all three boxes and  
 494 in the two models. In particular, at the very high wavenumbers, the forward cascade is  
 495 underestimated in the flux computed from the geostrophic velocity. A possible reason for

496 this mismatch at smaller scales could be explained by the findings of Brüggemann and Eden  
 497 (2015) that showed that ageostrophic flows at fine scales are an excellent catalyst for energy  
 498 cascade towards dissipation. Despite the differences in terms of flux magnitude, the overall  
 499 shape of the flux is consistent for the two forms of spectral flux. The scale at which the  
 500 inverse cascade is maximum is the same irrespective of the type of velocity fields.

## 501 **5 Discussion and Summary**

502 In this study, we presented the analysis of kinetic energy wavenumber spectral density,  
 503 slope, and flux by using datasets from daily and hourly outputs of two submesoscale permit-  
 504 ting ocean models of the North Atlantic. The analysis presented has shown that in summer,  
 505 the North Atlantic ocean follows the QG framework (with  $\sim k^{-3}$  spectral shape) and in  
 506 winter, the basin mostly reflects a  $k^{-2}$  spectral shape, a characteristic of a front dominated  
 507 regime. The estimated kinetic energy spectral flux revealed an overall net inverse cascade  
 508 of energy with a significant direct cascade of energy at high wavenumbers. The spectral  
 509 flux undergoes a seasonality that is associated with a stronger forward cascade at high  
 510 wavenumbers in winter. This increased forward cascade in winter is further amplified in the  
 511 presence of high-frequency motions. The spectral flux as a function of depth reveals that  
 512 the forward cascade at high wavenumbers is confined to the mixed layer while the inverse  
 513 cascade dominates the water column down to 700m. We showed that the maximum inverse  
 514 cascade occurs at a scale that coincides with the energy-containing scale.

515 Not until recently, most basin/global scale simulations had their outputs stored in  
 516 the form of daily averages. Our results show that high-frequency motions (that are only  
 517 resolved in hourly outputs) affect the distribution and exchanges of kinetic energy. We  
 518 observed that the difference between the daily and hourly results is mostly in the form of  
 519 an increased variance and (forward) cascade at fine-scales in favor of hourly fields. High-  
 520 frequency motions are dominated mainly by ageostrophic motions that include unbalanced  
 521 submesoscales and fast propagating internal gravity waves. These two classes of motions  
 522 are out of phase seasonally with stronger submesoscales in winter and stronger internal  
 523 gravity waves in summer. It is puzzling that while the exchanges of energy at fine-scale are  
 524 unaffected by high-frequency motions in summertime, the distribution of energy shows the  
 525 contrary. We observed higher variance at fine-scales in power spectral density estimated  
 526 from hourly fields. The reason for this disparity between the impact of high-frequency on  
 527 spectral density and flux in summertime is not apparent but would be an interesting subject  
 528 to investigate further.

529 The two kilometeric simulations used in this study have similar horizontal grid spacing  
 530 but different numerics, sub-grid parameterization, and vertical resolution. In particular,



531 NATL60 has 300 z levels, while HYCOM50 has 32 hybrid layers. Despite these differences,  
532 the two simulations agree well on the overall dynamics of the North Atlantic. However,  
533 HYCOM50 is more energetic compared to NATL60 both at the surface and in the interior.  
534 We found the estimated cascade in HYCOM50 to be of higher magnitude compared to  
535 NATL60 for both direct and inverse cascade. The difference in energy levels between the  
536 two models could be due to the difference in length of spin-up or/and the choice of subgrid-  
537 scale parameterization. Initially, we thought that HYCOM50 having just 32 hybrid layers  
538 in the vertical, could lead to a more surface intensified energy cascade in HYCOM50 than  
539 in NATL60. Nevertheless, this is not the case, because across all scales and at depth,  
540 HYCOM50 seems to show stronger energetic compared to NATL60.

541 NATL60 and HYCOM50 are designed mainly to serve as an observational dataset for the  
542 anticipated SWOT mission. SWOT will provide sea surface height, and by using geostrophic  
543 approximation, we would obtain the geostrophic velocity, a requisite for computing cross-  
544 scale energy transfer. At fine-scale, geostrophic approximation is however less accurate.  
545 Thus, accurately diagnosing surface velocity from sea surface height at fine-scales remains  
546 a challenge, and this has an impact on the estimate of kinetic energy cascade. Our results  
547 show that at fine-scales, not accounting for the ageostrophic motions could affect at all  
548 scales, the true magnitude of the estimated cascade of kinetic energy.

549 The results presented in this study are based on the output of ocean numerical sim-  
550 ulations that are forced with realistic atmospheric winds. Recent literature suggests that  
551 air-sea coupling at fine-scales could affect the evolution and energetics of oceanic eddies.  
552 Renault et al. (2016) using a coupled/uncoupled model of the California Upwelling System  
553 argued that the ocean-atmosphere interactions have feedback that acts as an oceanic eddy  
554 killer. This feedback deflects energy from the geostrophic current into the atmosphere and  
555 dampens geostrophic kinetic energy. A possible future study would be to recompute the  
556 (kinetic energy transfer) diagnostics in this study using datasets from an ocean-atmosphere  
557 coupled simulation. This sort of analysis would take into account the direct impact of air-sea  
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573 [data.html](http://meom-group.github.io/swot-natl60/access-data.html), respectively.

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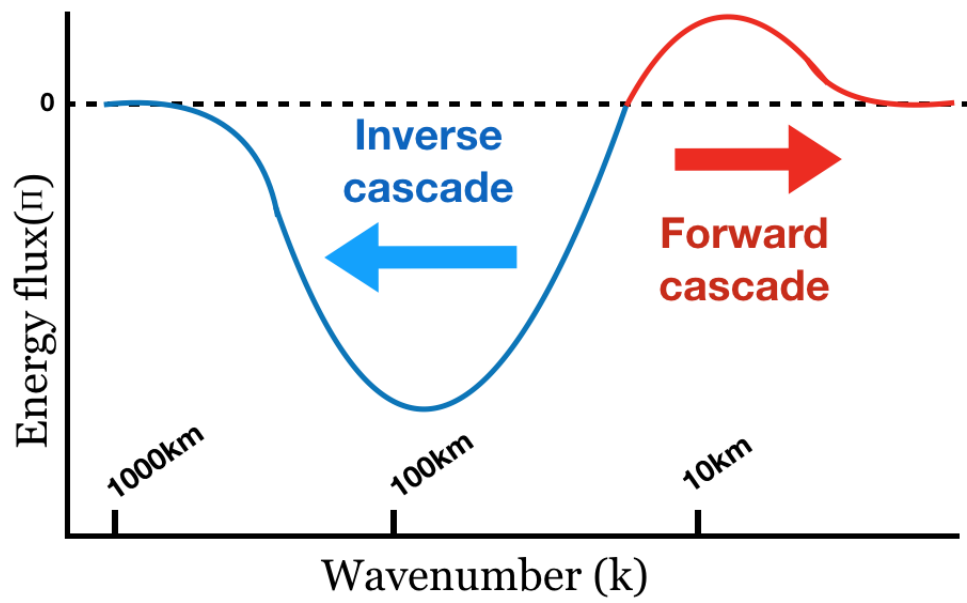
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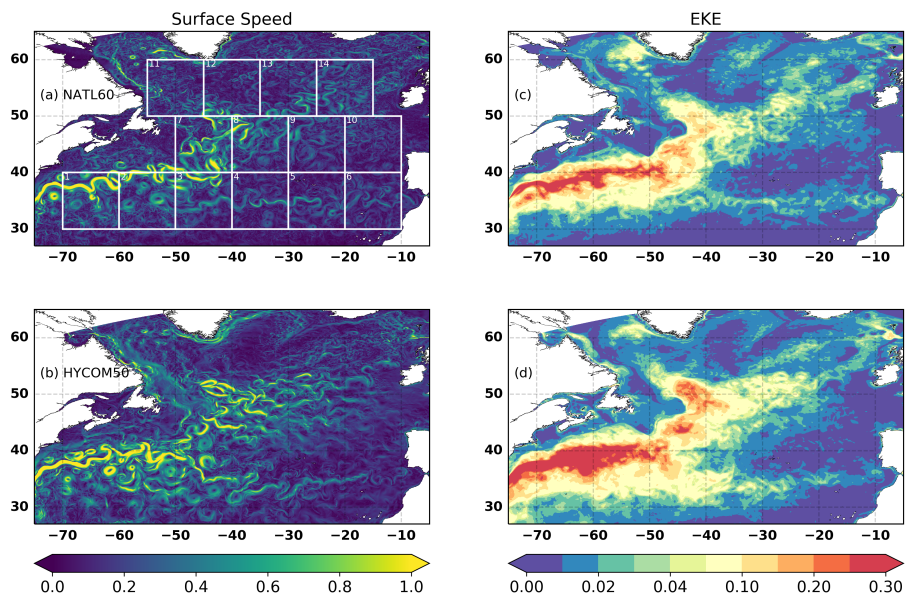
**Table 1.** Table of model parameters for NATL60 and HYCOM50

|                      | NATL60                         | HYCOM50                |
|----------------------|--------------------------------|------------------------|
| Domain               | 26.5N - 65N                    | 28 - 80N               |
| Numerical Code       | Nemo v.3.6                     | HYCOM                  |
| Horizontal grid      | 1/60: 0.9-1.6 km               | 1/50:1.1-2.2 km        |
| Vertical coordinate  | Z partial cells                | Hybrid (Z & isopycnal) |
| Integration period   | 6 Months                       | 20 years               |
| Vertical grid        | 300 Levels : 1-50 m            | 32 Layers              |
| Boundary conditions  | GLORYS2v3                      | GDEM                   |
| Atmospheric forcing  | DFS5.2                         | ERA-40                 |
| Horizontal Viscosity | Implicit in momentum advection | Laplacian & Biharmonic |

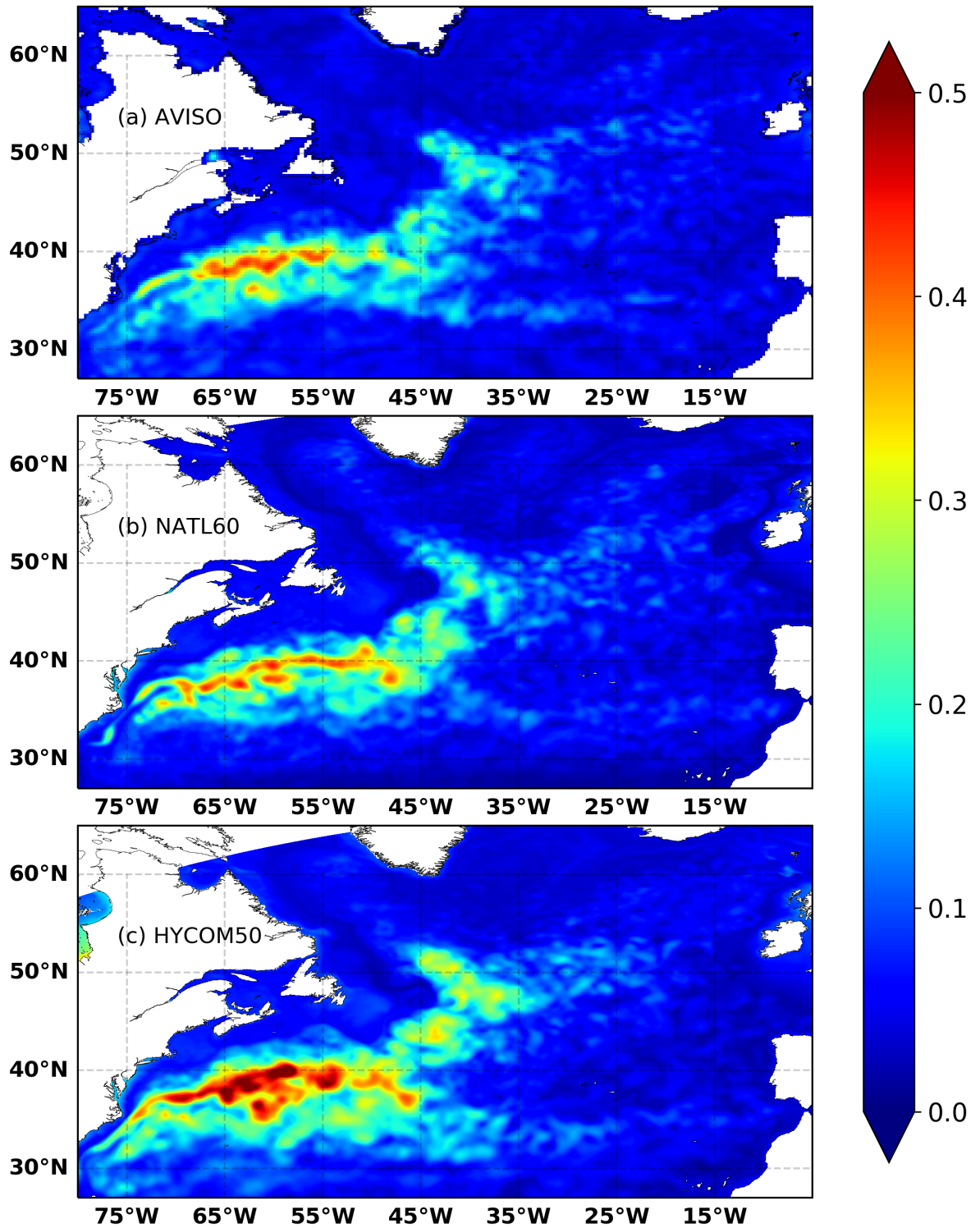


**Figure 1.** Schematics of kinetic energy spectral flux in the ocean at mid-latitude. Blue : inverse cascade of energy, Red : forward cascade of energy.

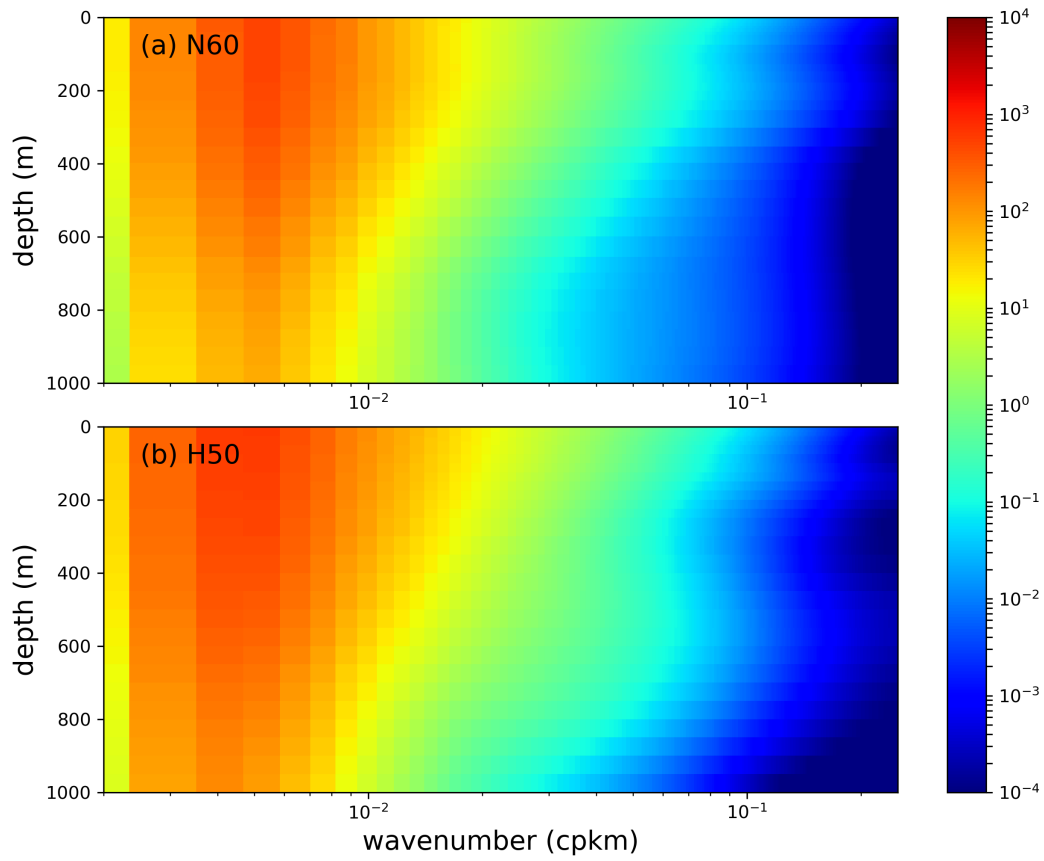




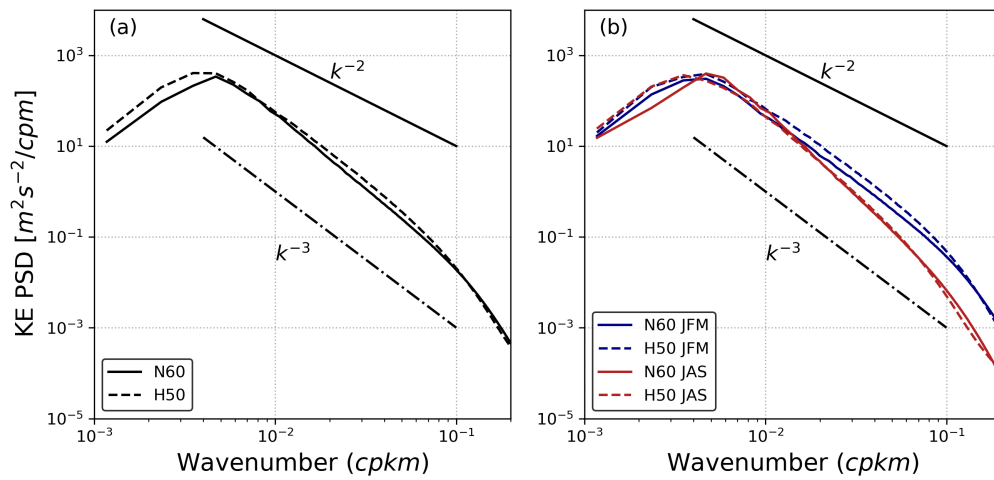
**Figure 2.** Left panel : snapshot of surface currents speed ( $m/s$ ) on march 1st for NATL60 (a) and HYCOM50 (b). Right panel : surface eddy kinetic energy ( $cm^2s^{-2}$ ) computed from daily output for NATL60 (c) and HYCOM50 (d).



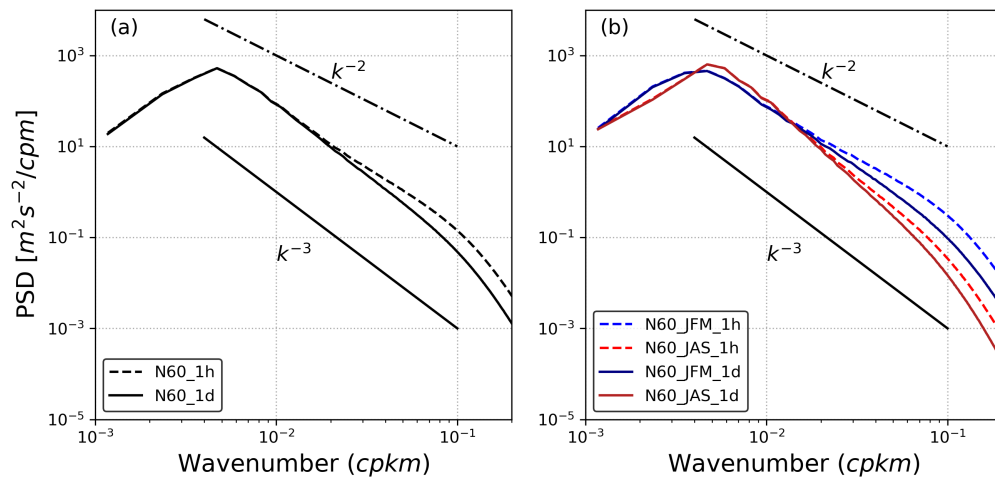
**Figure 3.** Standard deviation of sea surface height based on one year datasets for (a) AVISO, (b) NATL60 and (c) HYCOM50. The SSH values for NATL60 and HYCOM50 were degraded to 0.25° spatial resolution. In this comparison, we have used AVISO mean dynamical topography dataset from October 2012 to September 2013.



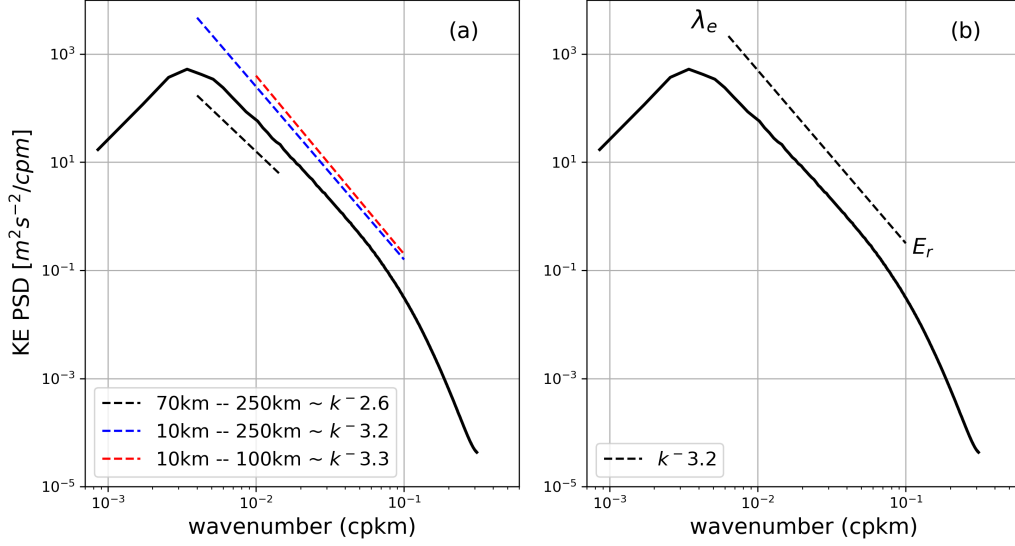
**Figure 4.** One year average of kinetic energy spectral density ( $m^2 s^{-2} / cpkm$ ) for Box 8 computed from horizontal total velocity as a function of depth for (a) NATL60 and (b) HYCOM50.



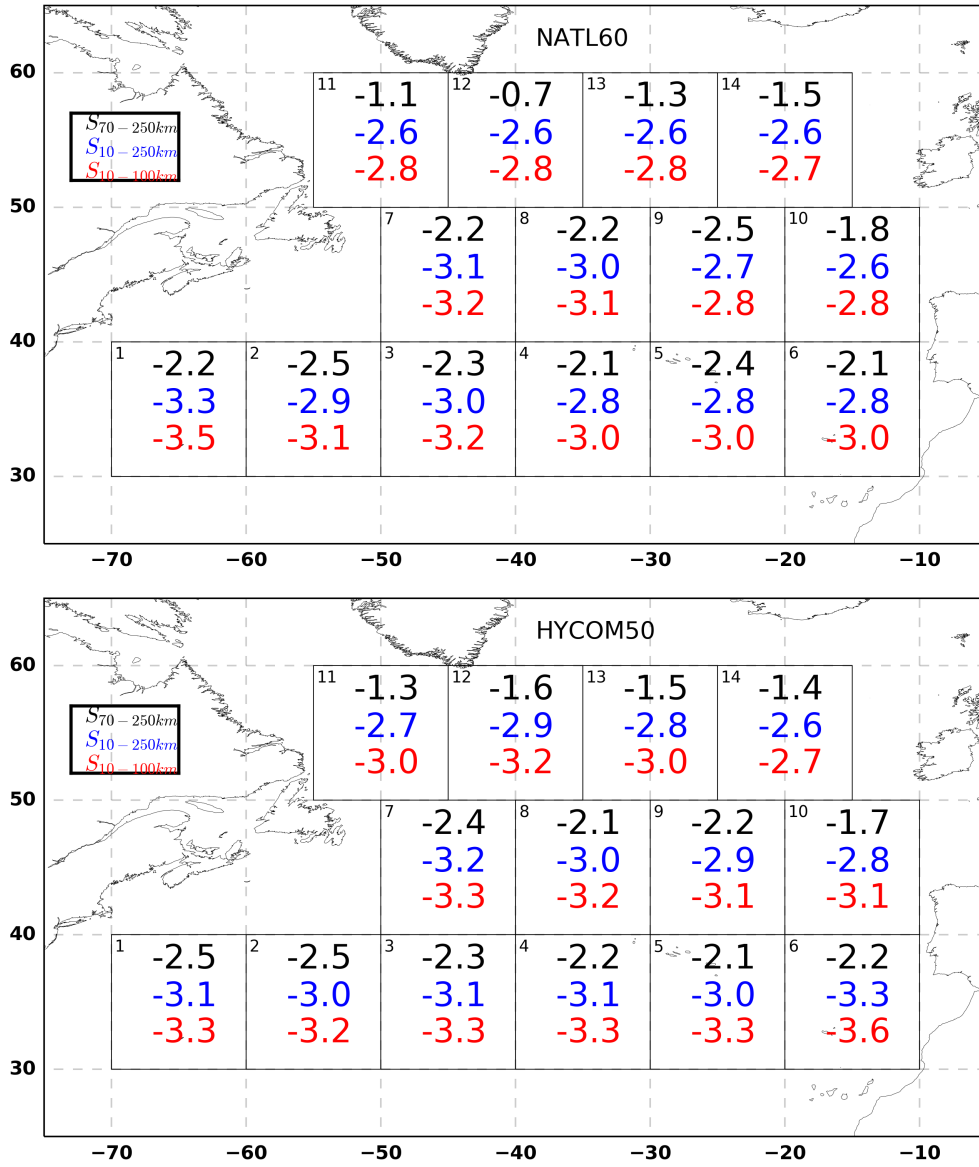
**Figure 5.** Kinetic energy spectral density for Box 8 (averaged over 1000 m depth) computed from daily output of horizontal total velocity for NATL60 (thick line) and HYCOM50 (dash line). (a) one year mean (b) winter (blue line) and summer (red line) averages. See Figure 5 in SI for a comparison of the surface vs depth averaged spectral density. A comparison of the spectral density between three different depth levels and the surface is presented in Figure 6 of SI



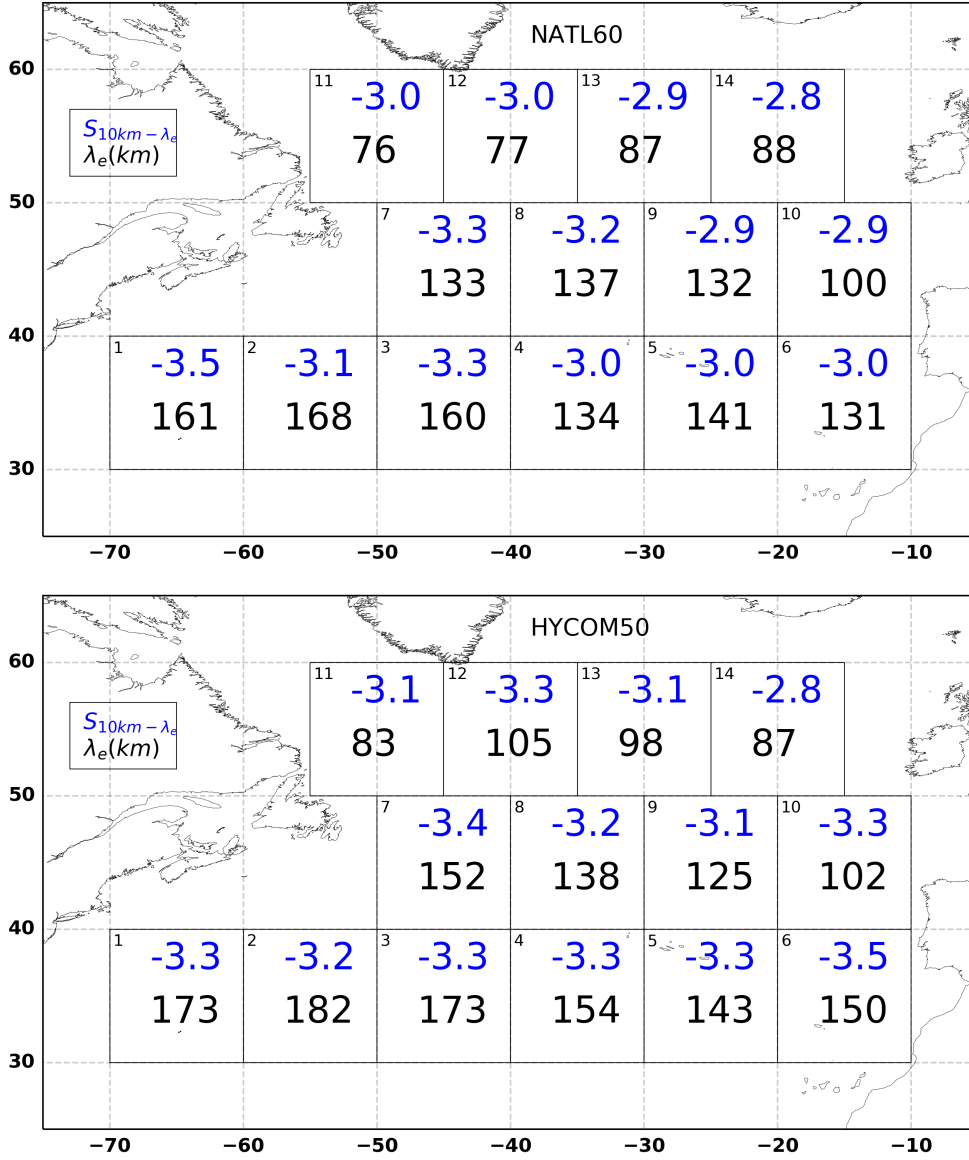
**Figure 6.** Comparison between surface kinetic energy spectral density computed from daily averages (thick line) and hourly averages (dash line) of velocity outputs for Box 8. (a) one year mean (b) winter (blue line) and summer (red line) averages. See Figure 7 and 8 in SI for a similar plots for all the boxes.)



**Figure 7.** (a) Average surface KE spectral density and slope for box 3 (NATL60) in the month of March for three different selected wavenumber ranges. The wavelengths ranges are represented by dashed lines with the color red, blue and black for 10-100 km, 10-250 km and 70-250 km respectively. (b) A schematic to illustrate the proposed dynamical approach to estimate spectral slope.  $\lambda_e$  is the energy-containing scale (which represents the scale of the most energetic eddy structure) and it is estimated from the kinetic energy wavenumber spectral density by using equation (3) while  $E_r$  is the effective resolution (a function of the model grid-size) and is taken as  $5 \times$  the model grid size.  $E_r$  is roughly equally to 10 km for both NATL60 and HYCOM50.

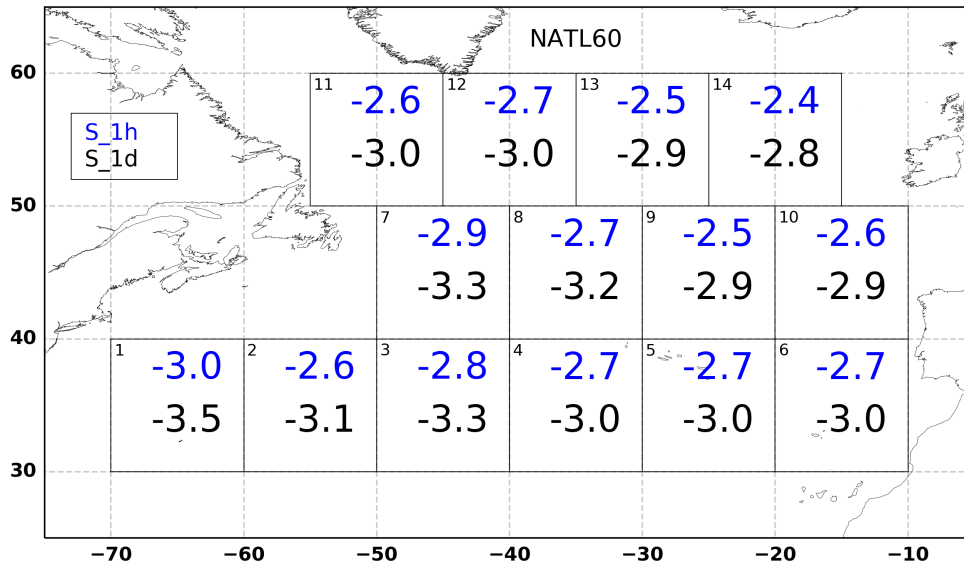


**Figure 8.** Map of spectral slope estimated from the surface kinetic energy spectral density for three different selected wavenumber ranges. Colour red, blue and black represent 10 - 100 km , 10 - 250 km and 70 - 250 km respectively

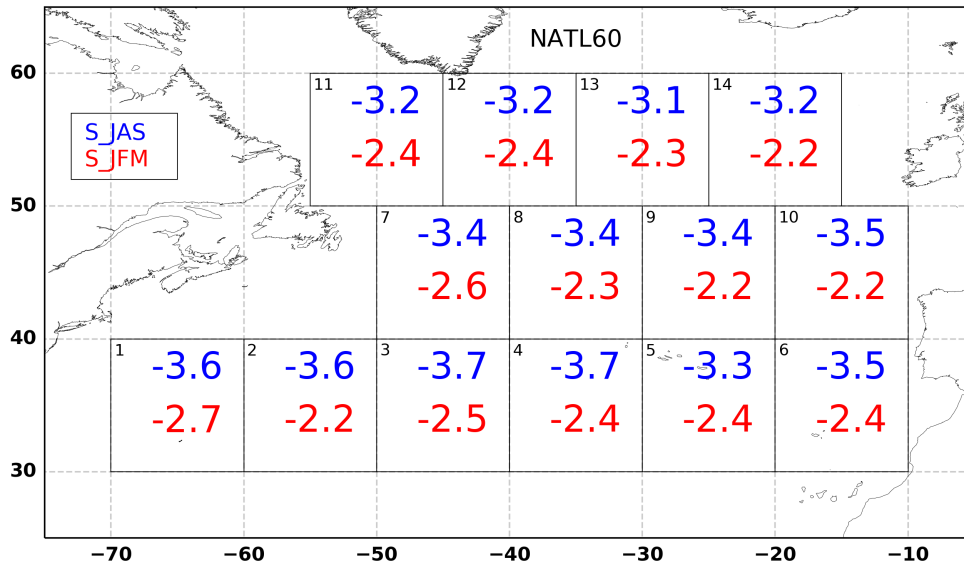


**Figure 9.** Map of spectral slope (blue colour) and energy containing scale,  $\lambda_e$  (black colour) from the surface kinetic energy spectral density. The slope is estimated between the model effective resolution ( $E_f$ ) and the energy containing scale ( $\lambda_e$ ).

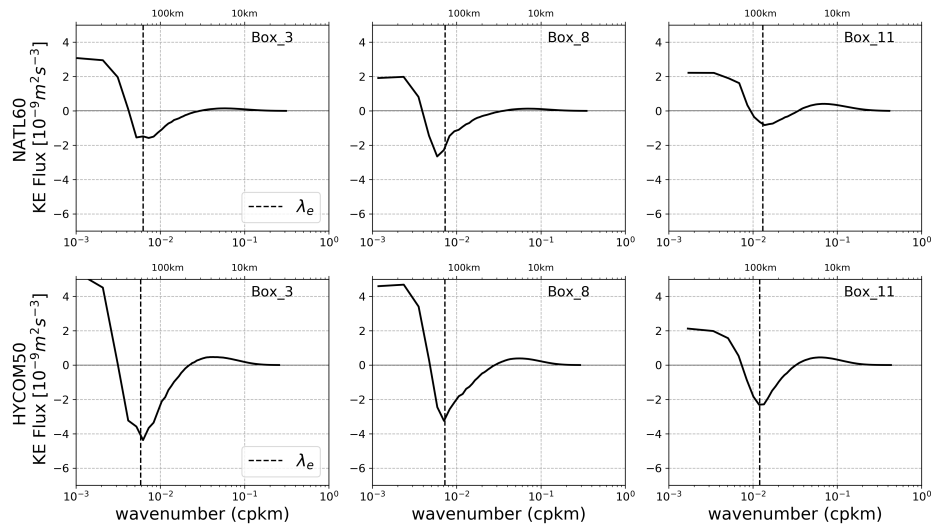




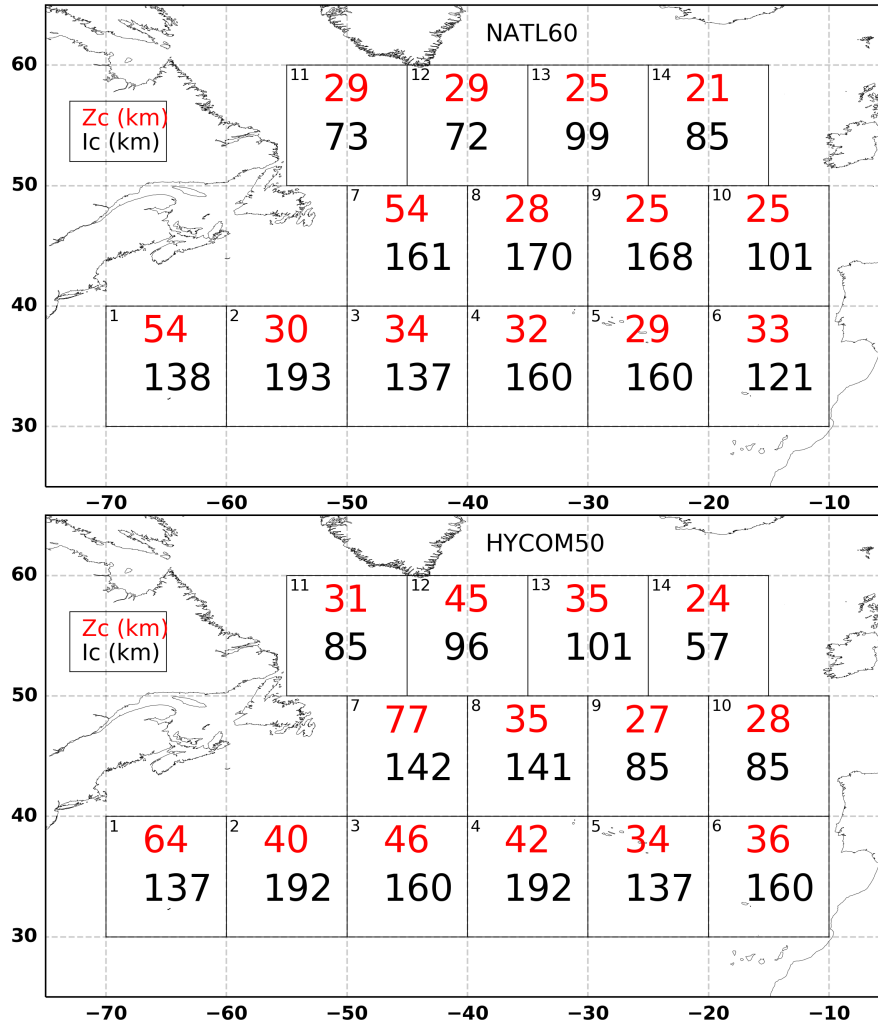
**Figure 10.** Map of spectral slope from the surface kinetic energy spectral density computed from daily versus hourly Fields. The slope is estimated between the model effective resolution ( $E_f$ ) and the energy containing scale ( $\lambda_e$ ).



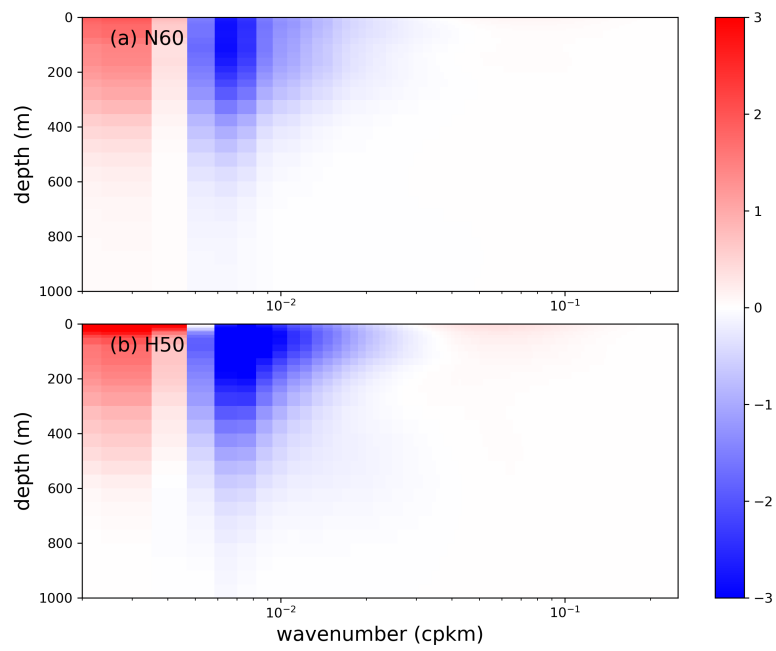
**Figure 11.** Map of spectral slope from NATL60 hourly surface kinetic energy spectral density in winter (JFM) and summer (JAS). The slope is estimated between the model effective resolution ( $E_f$ ) and the energy containing scale ( $\lambda_e$ ).



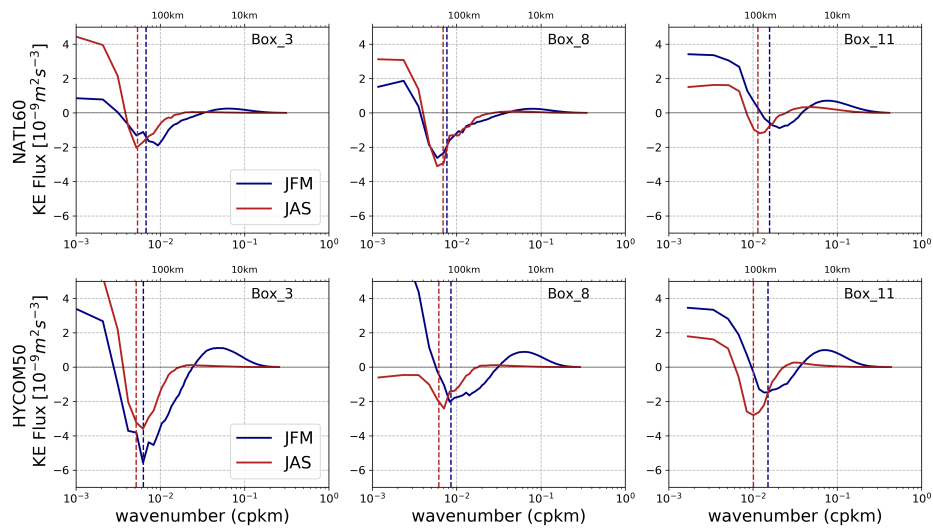
**Figure 12.** One year average of surface kinetic energy spectral flux computed from the daily output of horizontal total velocities. NATL60 (upper panel) and HYCOM50 (lower panel)



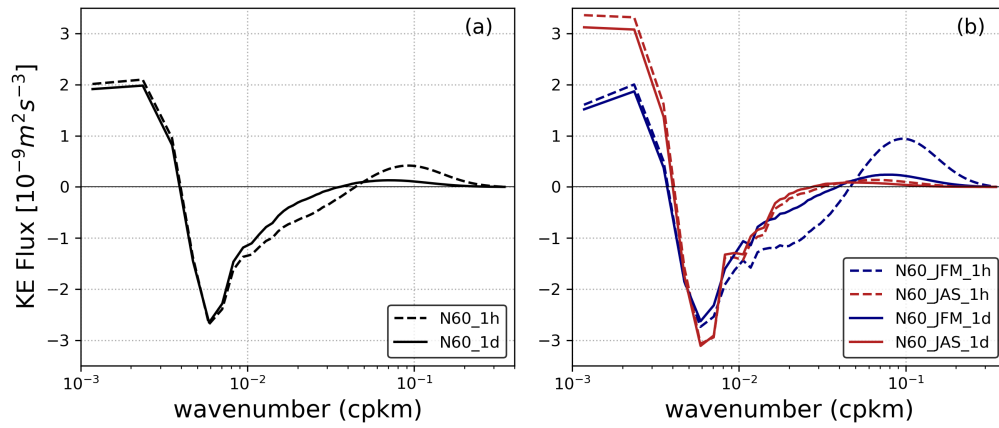
**Figure 13.** A geographical map of the scale of the kinetic energy spectral flux zero-crossing (red) and the most intense inverse cascade (black) estimated from one year average of the surface kinetic energy spectral flux for (a) NATL60 and (b) HYCOM50.



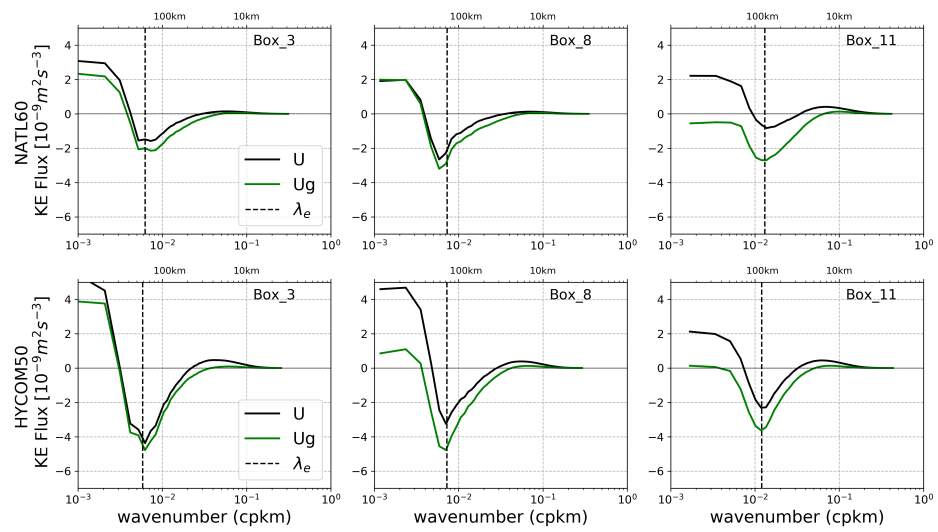
**Figure 14.** One year average of kinetic energy spectral flux for Box 8 computed from horizontal total velocity as a function of depth for (a) NATL60 and (b) HYCOM50.



**Figure 15.** Winter (blue line) and summer (red line) average of surface kinetic energy spectral flux computed from daily output of horizontal total velocities. Dash lines represents the energy containing scale. NATL60 (upper panel) and HYCOM50 (lower panel)



**Figure 16.** Comparison between surface kinetic energy spectral flux computed from daily averages (thick line) and hourly averages (dash line) of velocity outputs for Box 8. (a) one year mean (b) winter (blue line) and summer (red line) averages



**Figure 17.** Surface kinetic energy spectral flux computed from total velocity (black line) versus geostrophic velocity (green line). NATL60 (upper panel) and HYCOM50 (lower panel)